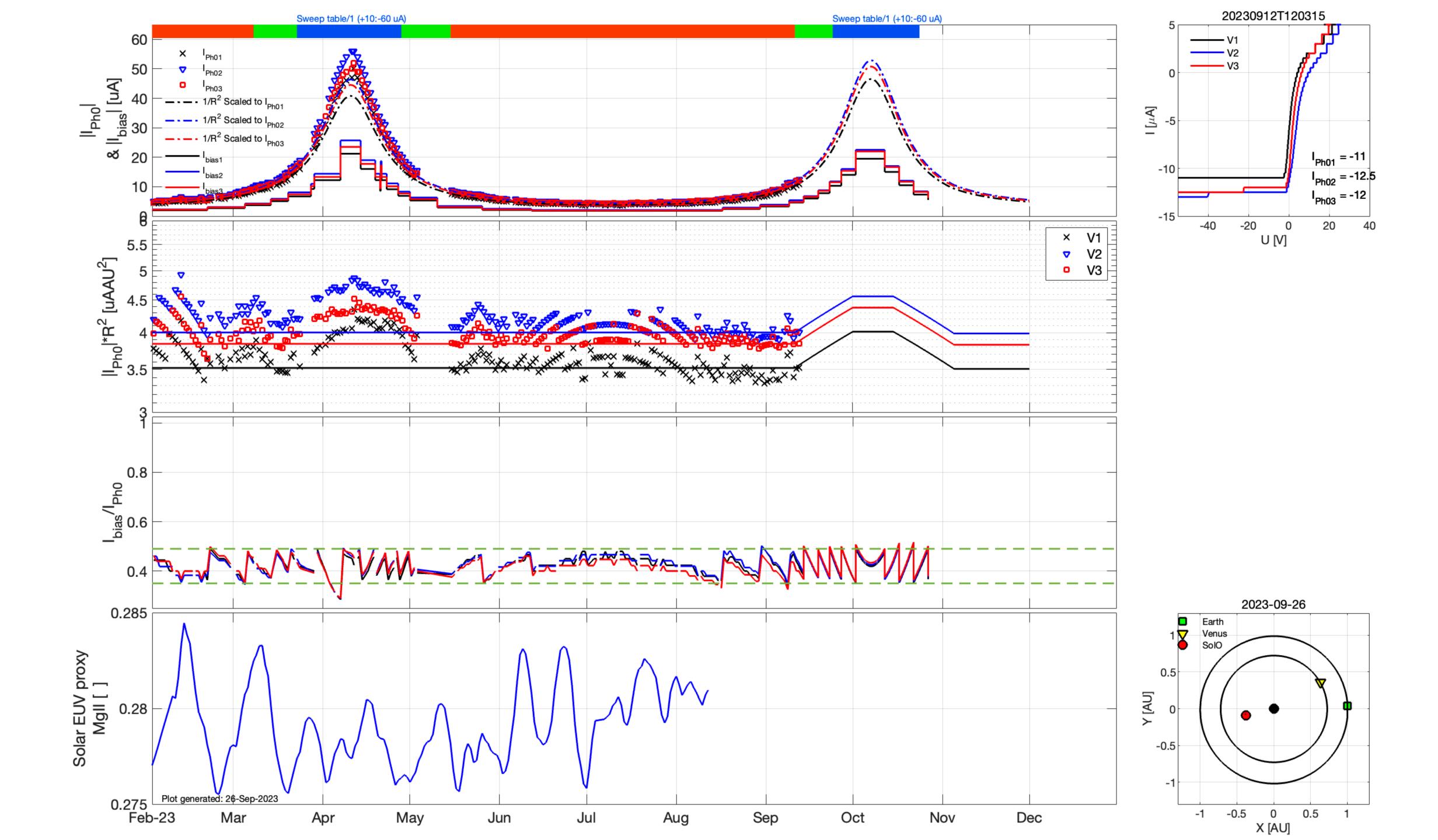


BIAS Status RPW Consortium meeting

Yuri Khotyaintsev, 2023-10-03



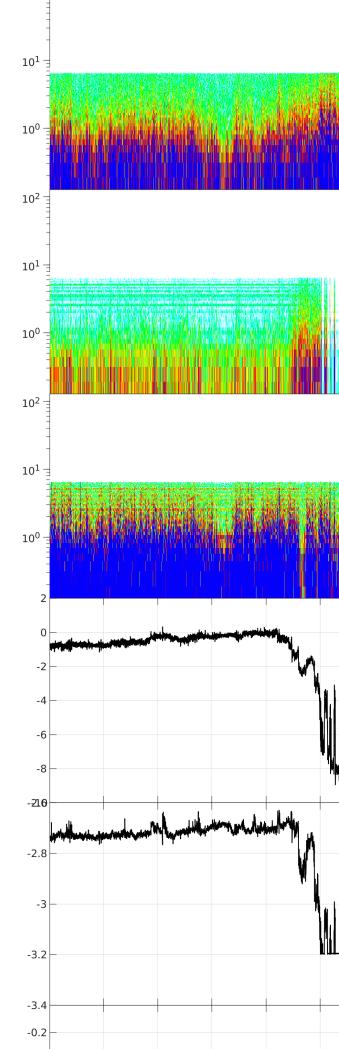
Saturation

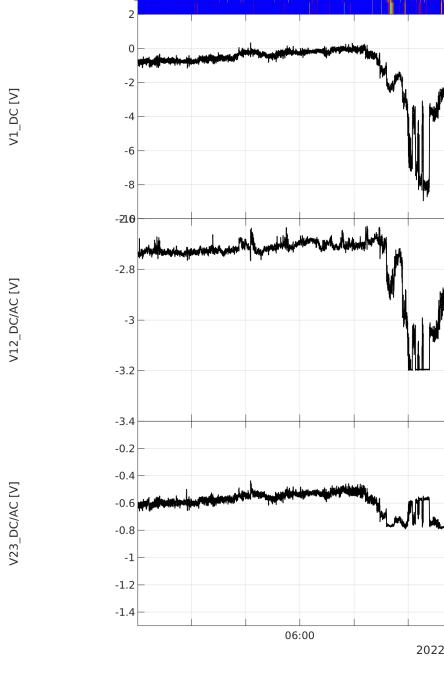


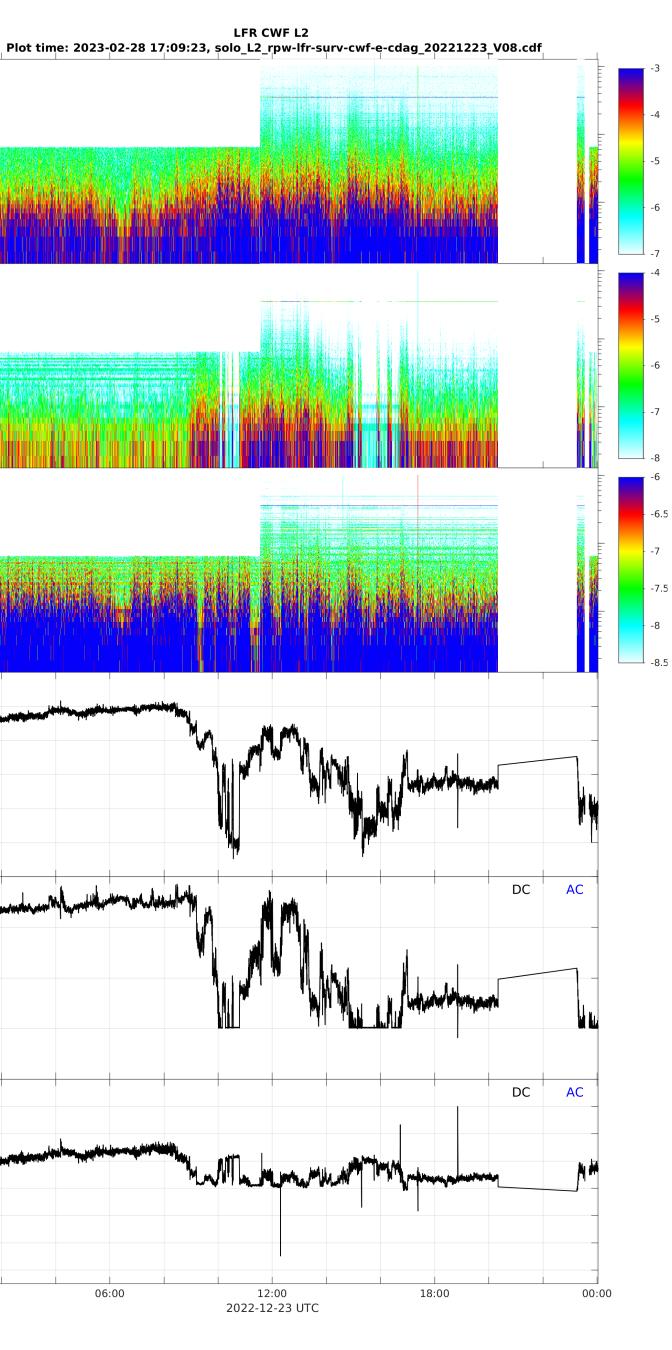
V12_DC/AC f [Hz]

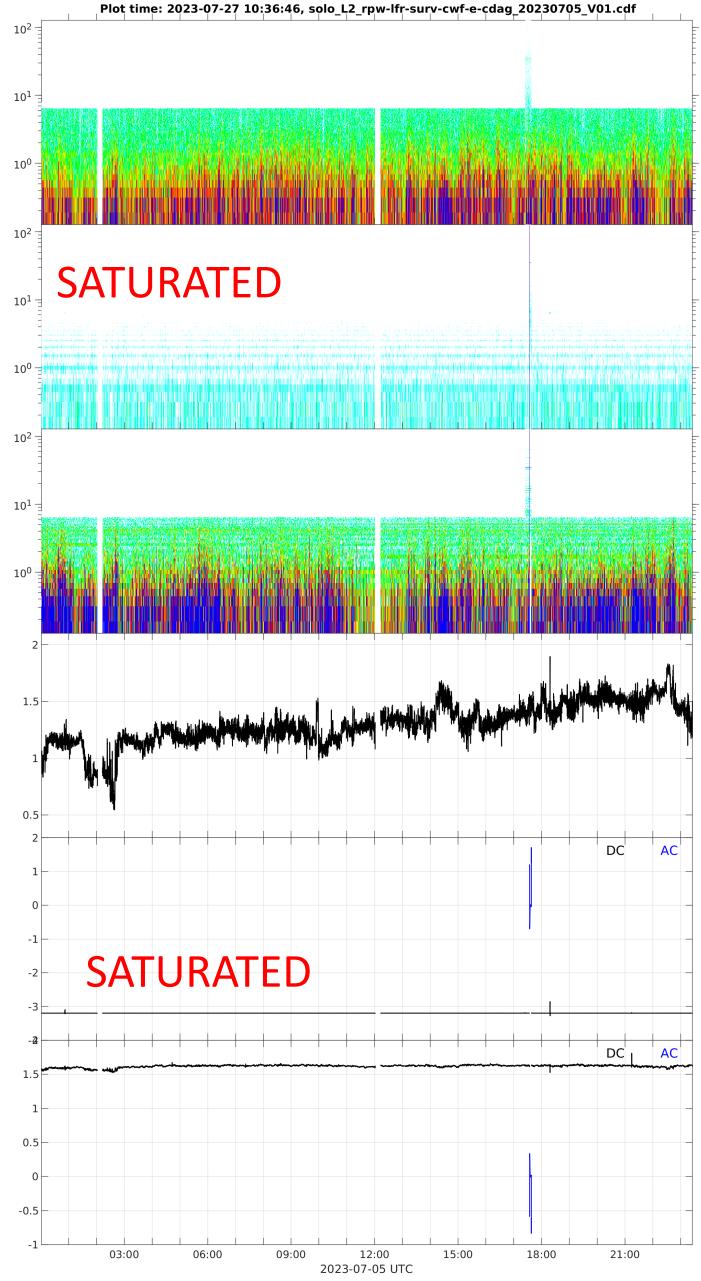
V23_DC/AC f [Hz]

10









LFR CWF L2

V1_DC f [Hz]

V12_DC/AC f [Hz]

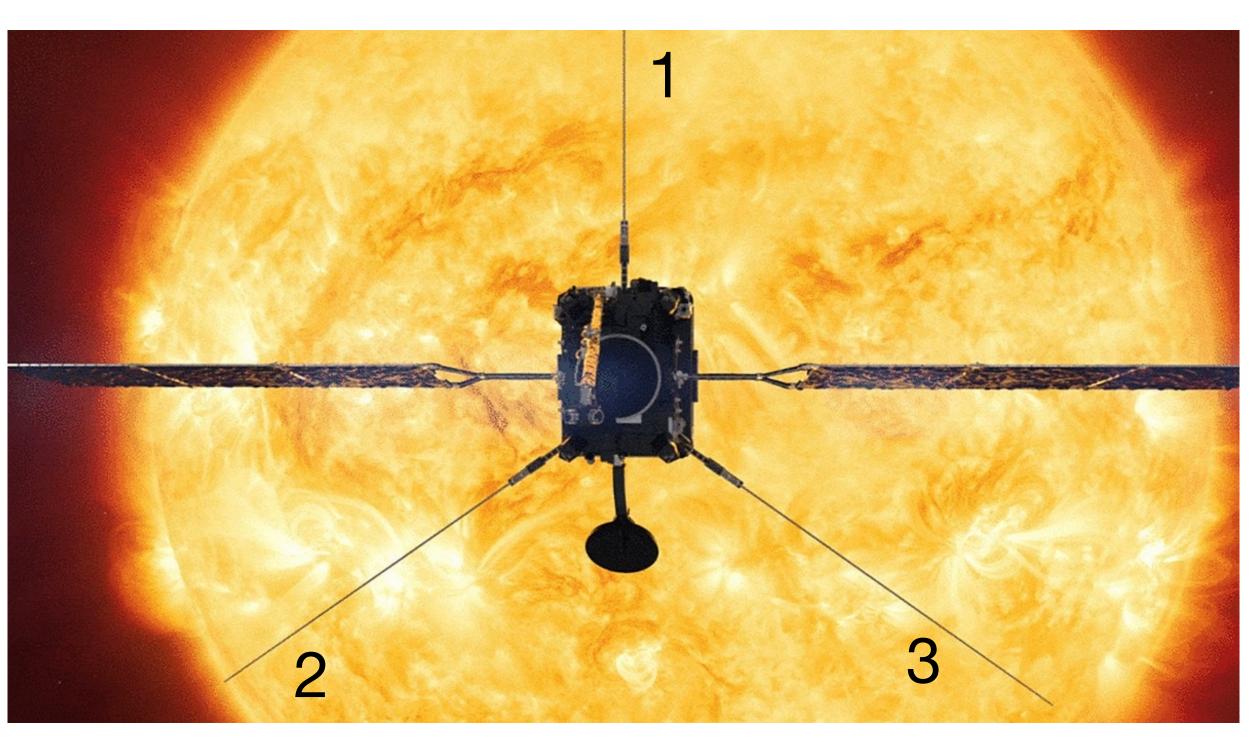
V23_DC/AC f [Hz]

V1_DC [V]

V12_DC/

V23_

-3	
4	-
5	og ₁₀ [V ² /Hz]
6	log
-7	
-4	
5	1
6	log ₁₀ [V ² /Hz]
7	log
-8	
6.	5
7	ы log ₁₀ [V ² /Hz]
7.	2 2 10
8	
-8.	5



The signals we normally measure: V1_DC V12_DC (V1-V2) V23_DC (V2-V3)

V2 & V3 are then computed from the diffs.

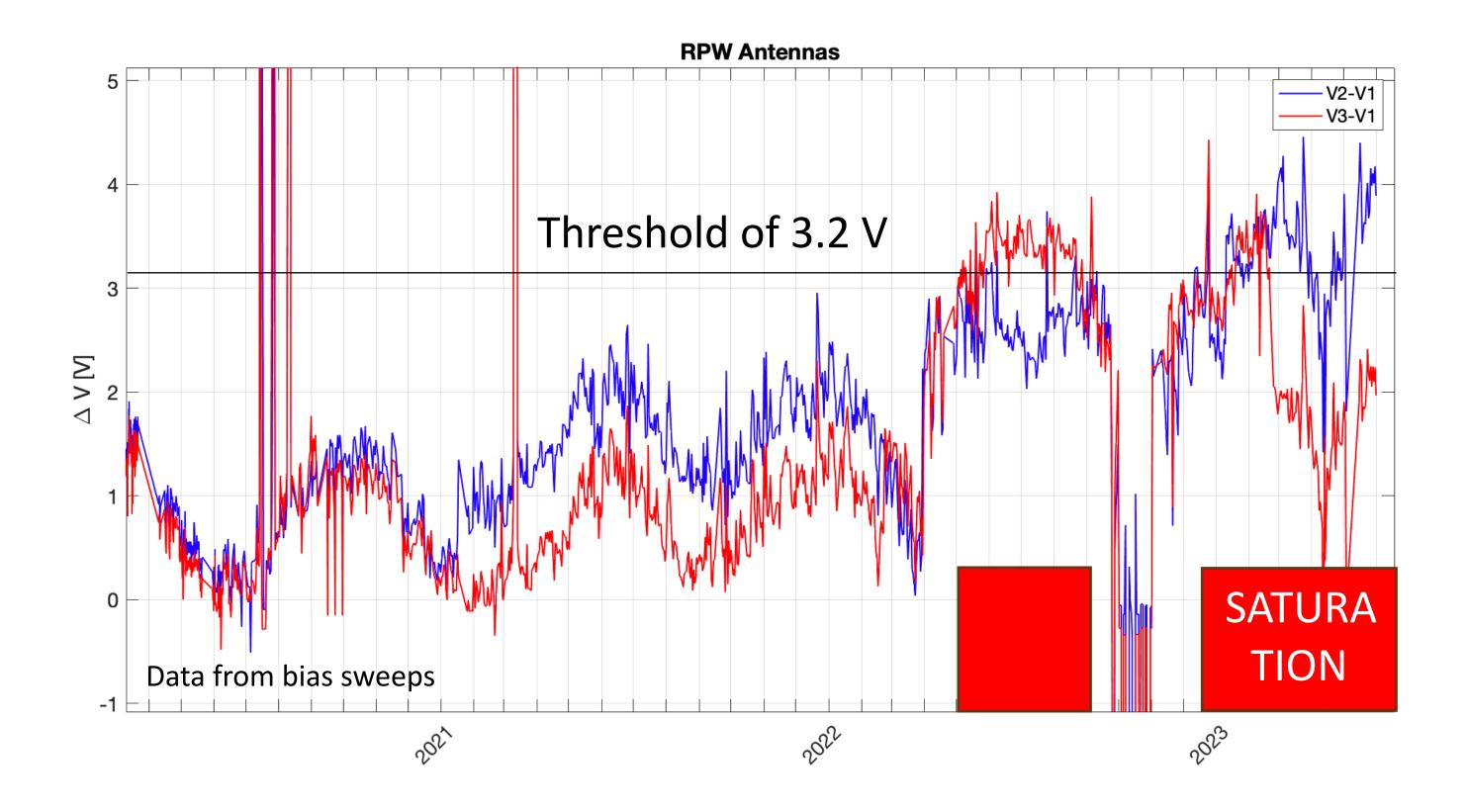
The analogue diffs are limited to \pm -3.2 V

- corresponds to E-field of ~500 mV/m,
- Solar wind VxB ~5 mV/m









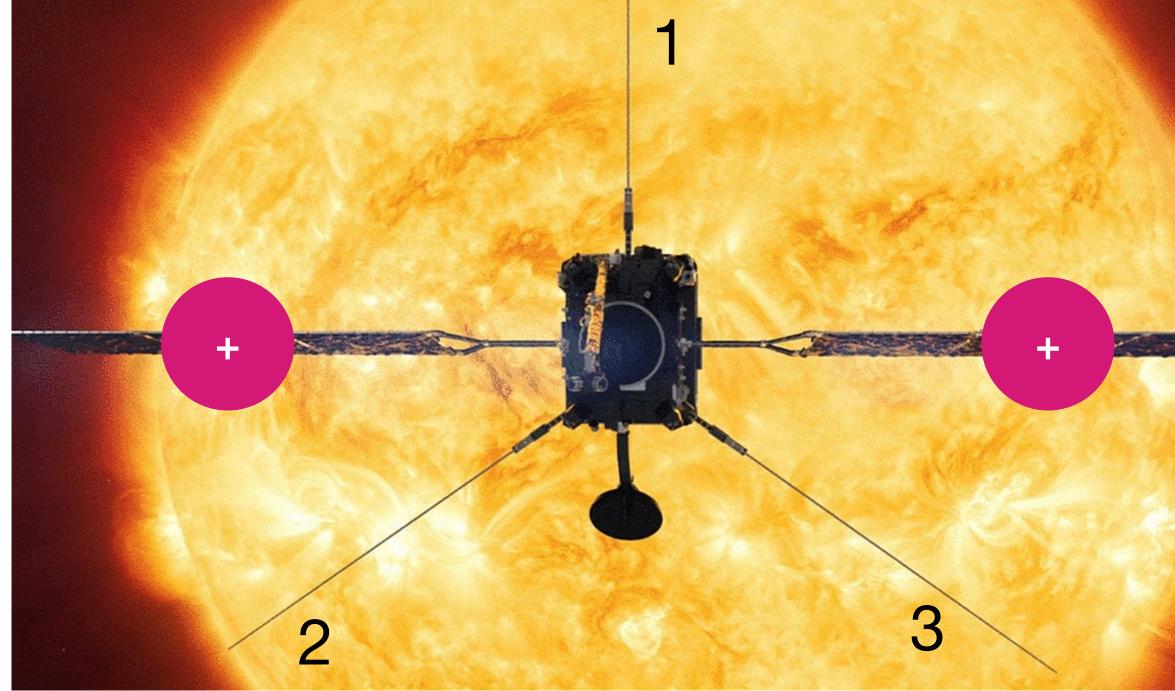
V12 was increasing over time and started to exceed 3.2 V in December 2022, leading to saturation of the DC E-field measurement. This implies that only the V23 component is available. Spacecraft potential and density can be computed from V1.

As a workaround, V12 was replaced by V13 on August 21, 2023. However, even V13 exceeded the saturation threshold in 2022 and early 2023, so we can reencounter the saturation.

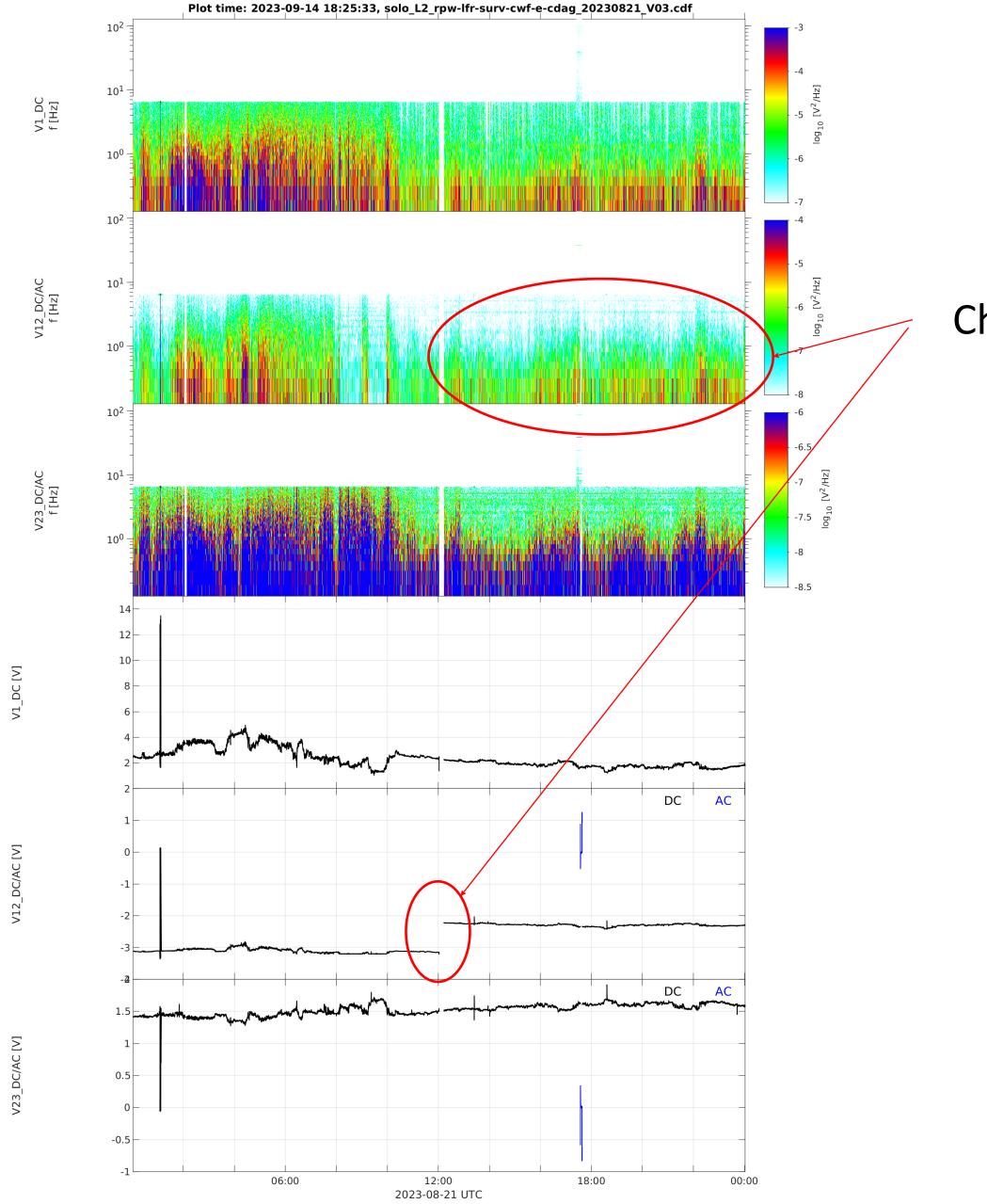
Evolution of saturation

The cause

- Micrometeorites hit the solar panels
- This blows fuses and parts of the panel become disconnected from the SC ground
- Non-grounded areas charge up
- Antennas 2 & 3 are close to the panels and see these charges



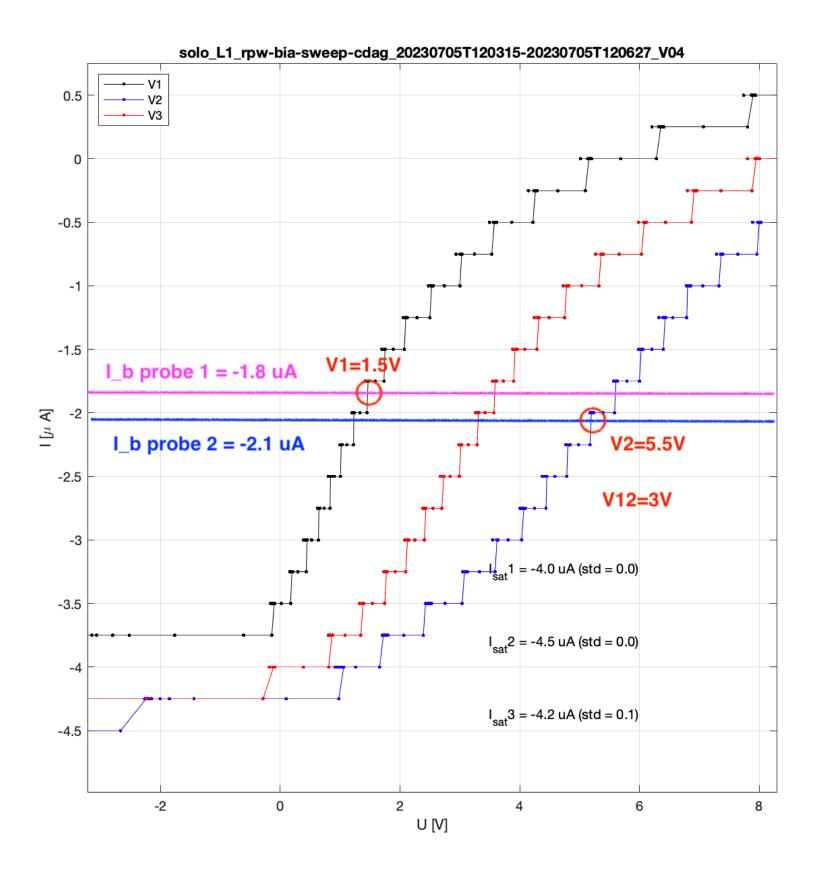




LFR CWF L2

Workaround

Change from V12 to V13. Data is good after this.



Saturation

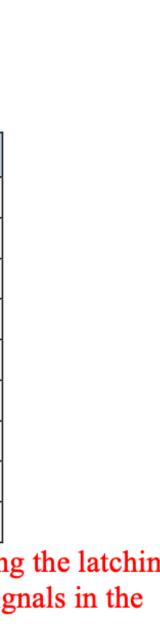
- Things already done
 - Switched to V13 instead of V12
 - Used V1 only to produce SCPOT
 - Implemented monitoring of saturation
- Things to be done \bullet
 - Develop a flag to mark the saturation in the EFIELD data.
 - Switch processing to use V1 and recalibrate/reprocess everything.
 - Be ready to switch back to V12 \bullet
 - resurrected.

	BIAS outp	out signals				
MODE	BIAS_1	BIAS_2	BIAS_3	BIAS_4	BIAS_5	Operation
0	V1_DC	V12_DC*	V23_DC	V12_AC*	V23_AC	Standard operation
1	V2_DC	V3_DC	V23_DC	V12_AC*	V23_AC	probe 1 fails
2	V1_DC	V3_DC	V13_DC*	V13_AC*	V23_AC	probe 2 fails
3	V1_DC	V2_DC	V12_DC*	V12_AC*	V23_AC	probe 3 fails
4	V1_DC	V2_DC	V3_DC	V12_AC*	V23_AC	Calibration mode 0
5	2.5V Ref	2.5V Ref	2.5V Ref	V12_AC*	V23_AC	Calibration mode 1
6	GND	GND	GND	V12_AC*	V23_AC	Calibration mode 2
7	GND	GND	GND	V12_AC*	V23_AC	Calibration mode 2

where * denotes the ability to switch from V12_AC and V12_DC, to V13_AC and V13_DC, using the latching relay, in the case of failure of probe 2. It is not possible to have a combination of V12 and V13 signals in the same mode.

Table 4 LFR/TDS BIAS unit output signals.

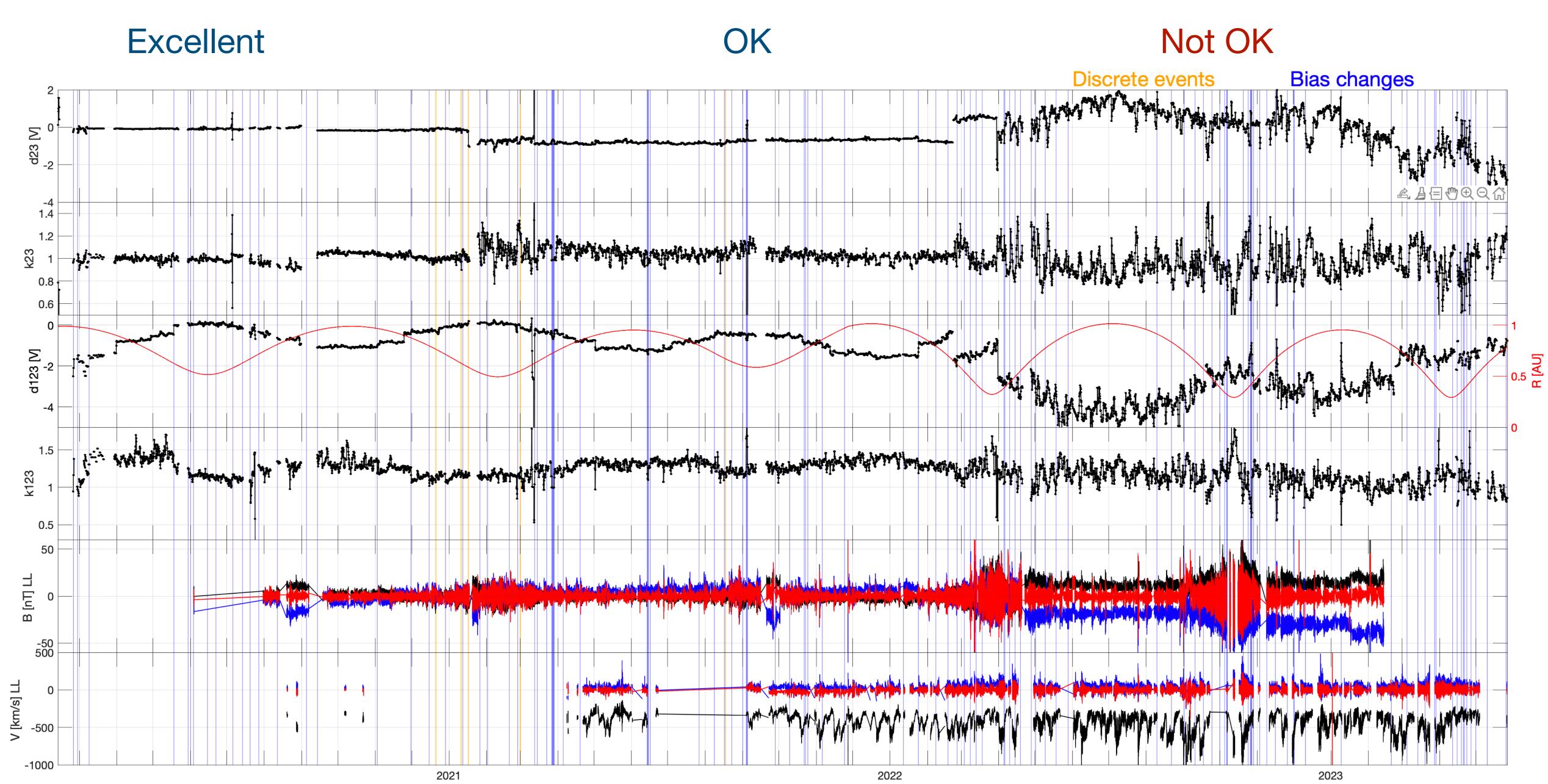
Be ready to switch to V1, V2, V3 (calibration mode 0) - procedure from commissioning needs to be



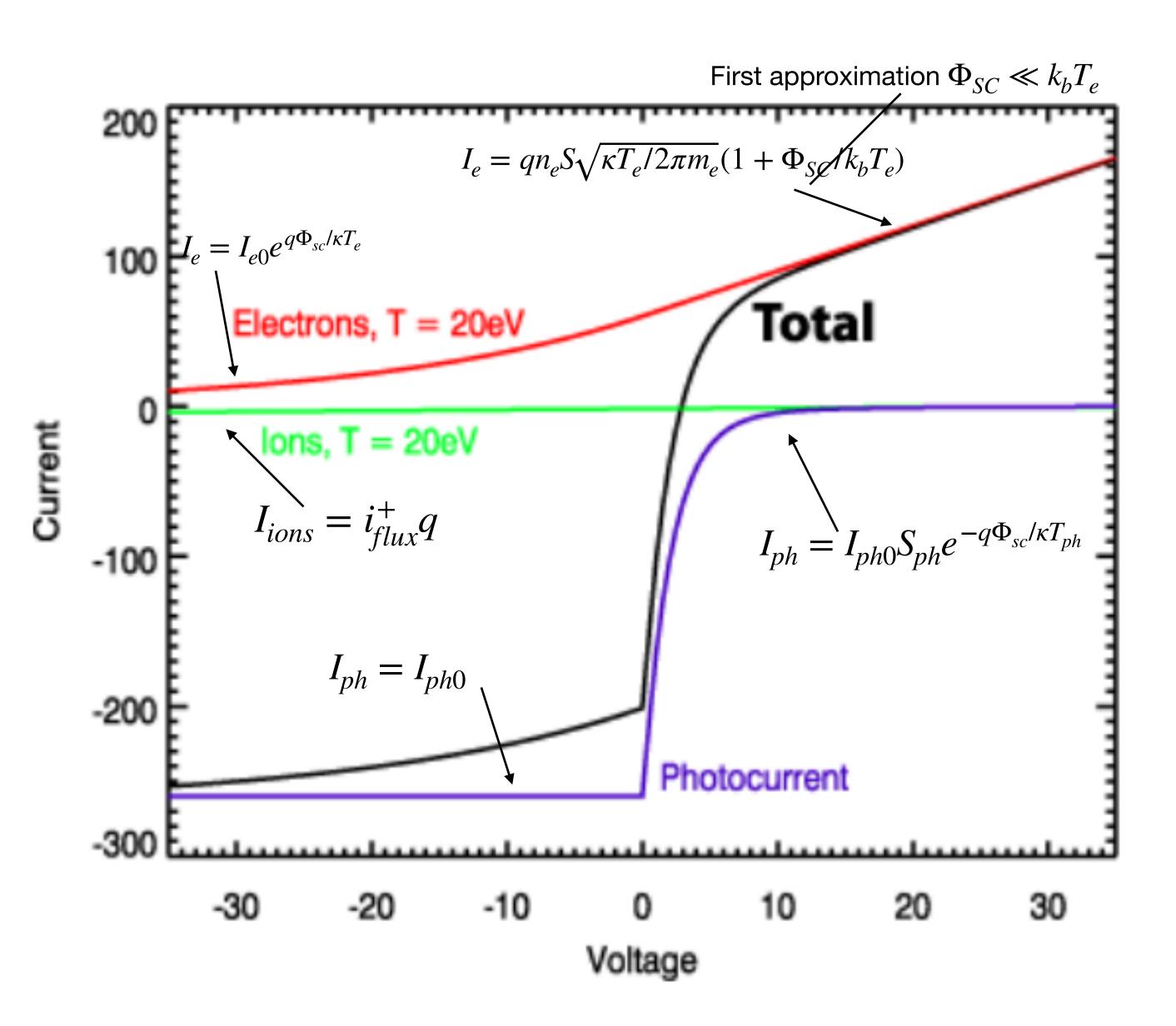
Status of datasets

- Density
 - Delivered up to 2022-11-25
 - Implemented a flag for high density (>100 cc)
 - Considering changing the ScPot to use V1 only.
 - Developing a new calibration approach (lead by Jordi Boldu)
- E-Field
 - Data in SOAR until Dec 1, 2022.
 - Preliminary cal until end of May 2023.

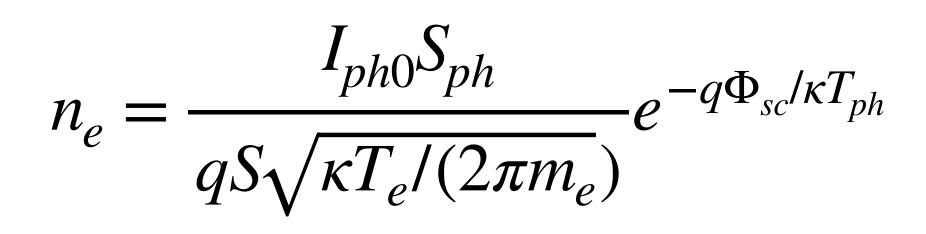
DC E-field

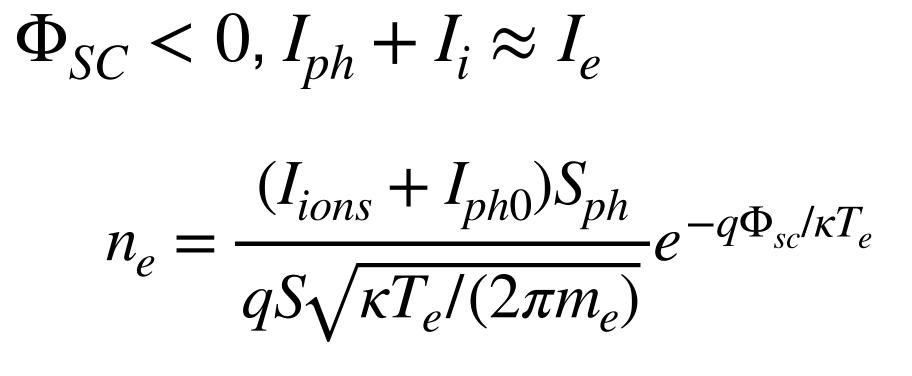


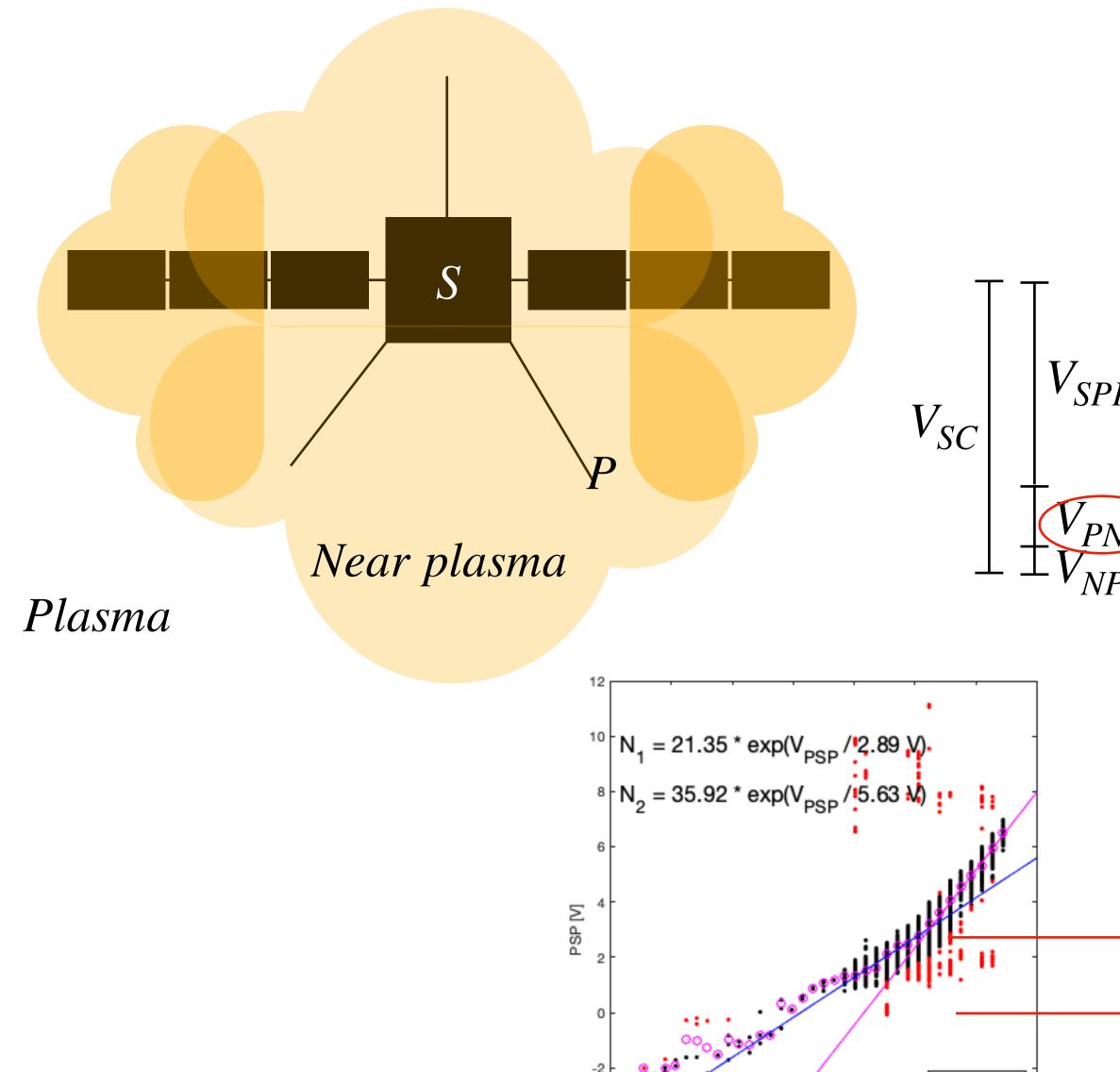
New density cal



 $\Phi_{SC} \ge 0, I_{ph} \approx I_e$







• data • median fit₁ fit₂
 outliers 2 2.5 3 3.5 4.5 5 4 log(Ne) [cc]

• • • •

-2

-6 1.5

$$n_e = n_0 e^{-\Phi_{sc}/\beta^{\pm}}$$

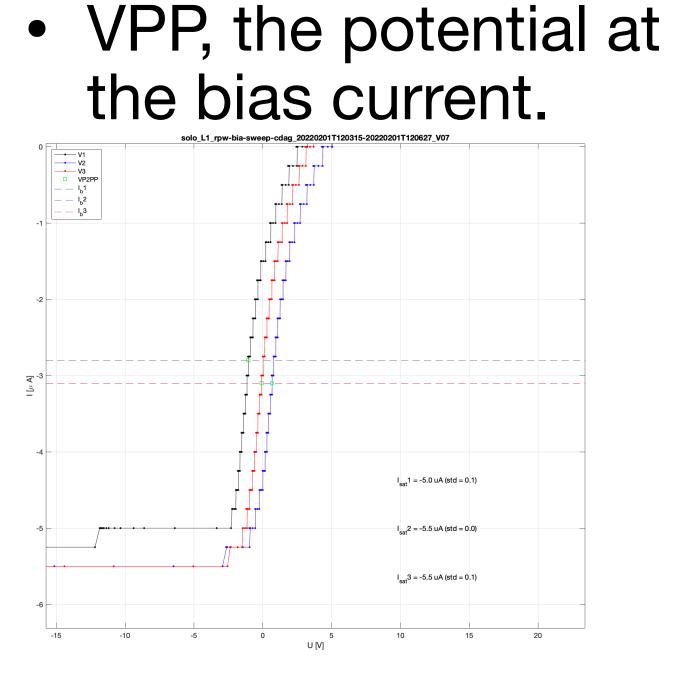
We have different expressions for
ne depending on the sign of the
potential. i.e at
$$\Phi_{sc} = 0$$

 $\Phi_{sc} = -\Phi_{PSP} + \Phi_P$
 $n_e = n_0 e^{(\Phi_{psp} - \Phi_{pp})/\beta}$
 $1 - V_{PP}$
 $ln(n_e) = ln(n_0) + \frac{\Phi_{PSP}}{\beta} - \frac{\Phi_P}{\beta}$

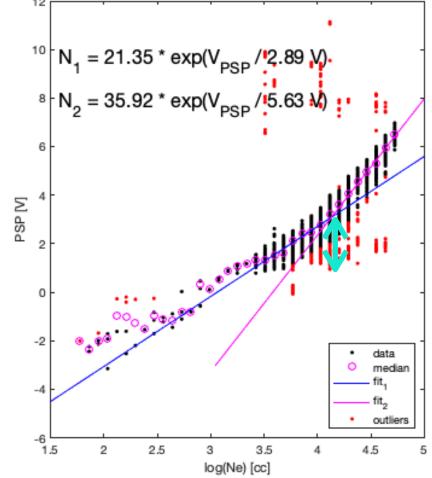


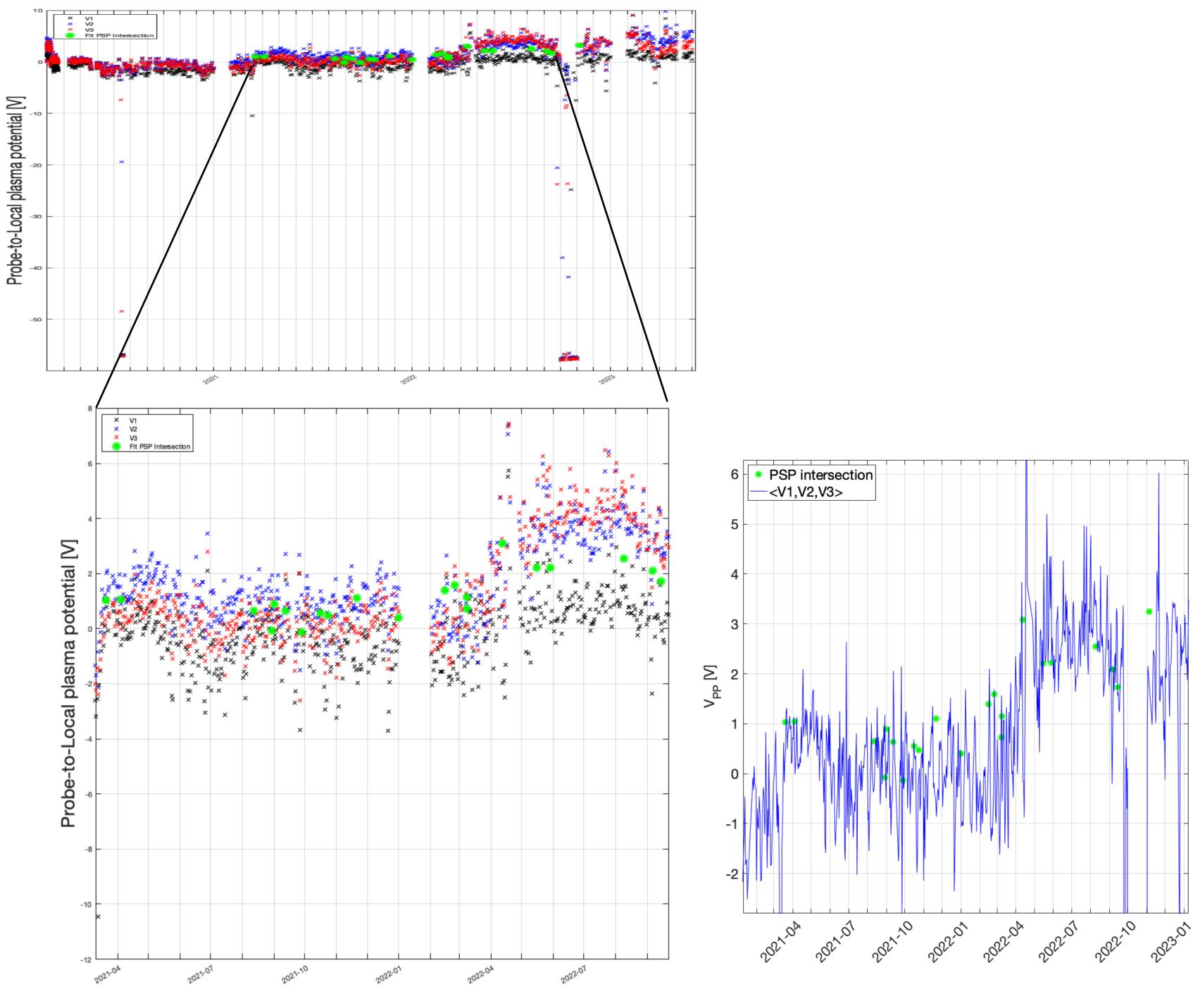


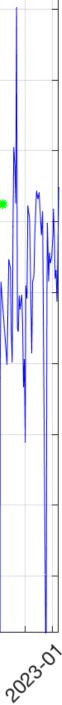




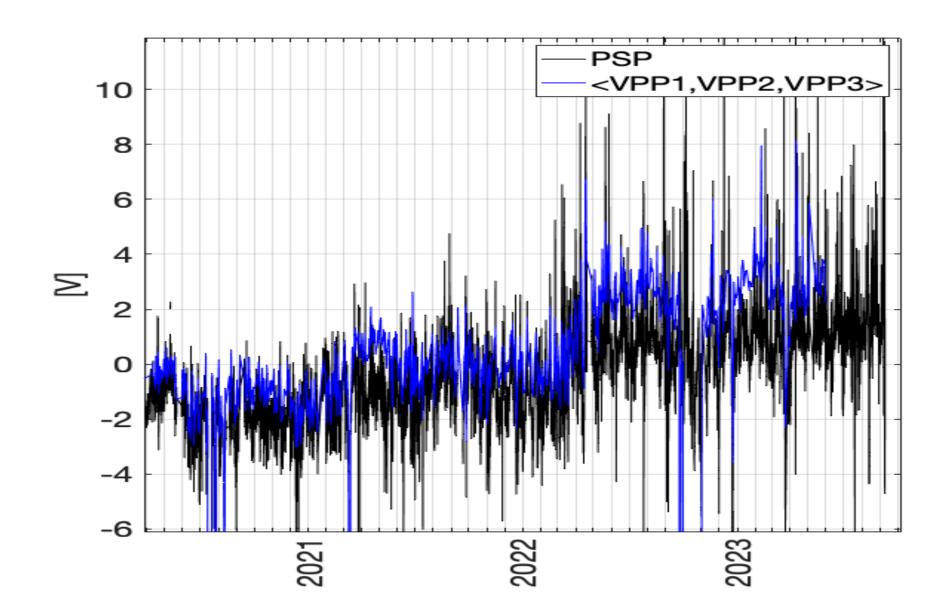
• PSP intersection: s/c potential value where the two fits intersect.







being close to 0.



• We start needing 2 fits around the time where PSP = Vpp, i.e. Vsc started

