



Solar Orbiter's first Venus flyby: Overview of RPW, MAG, and EPD in-situ observations and results from VGAM2 (:

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RPW, MAG and EPD teams



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RPW meeting
30/11/2021



SORBONNE université
UNIVERSITÉ PARIS-SACLAY



Solar Orbiter's first Venus flyby: observations from the Radio and Plasma Wave instrument

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Energetic Ions in the Venusian System: Insights from the First Solar Orbiter Flyby

R. C. Allen¹, I. Cernuda², D. Pacheco³, L. Berger³, Z. G. Xu³, J. L. Freiherr von Forstner³, J. Rodríguez-Pacheco², R. F. Wimmer-Schweingruber³, G. C. Ho¹, G. M. Mason¹, S. K. Vines¹, Y. Khotyaintsev⁴, T. Horbury⁵, M. Maksimovic⁶, L. Z. Hadid⁷, M. Volwerk⁸, A. Dimmock⁴, L. Sorriso-Valvo^{4,9}, K. Stergiopoulou⁴, G. B. Andrews¹, V. Angelini⁵, S. D. Bale^{10,11}, S. Boden^{3,12}, S. I. Böttcher³, T. Chust⁷, S. Eldrum³, P. P. Espada², F. Espinosa Lara², V. Evans⁵, R. Gómez-Herrero², J. R. Hayes¹, A. M. Hellín², A. Kollhoff³, V. Krasnoselskikh¹³, M. Kretzschmar^{13,14}, P. Kühl³, S. R. Kulkarni^{3,15}, W. J. Lees¹, E. Lorfèvre¹⁶, C. Martin^{3,17}, H. O'Brien⁵, D. Plettemeier¹⁸, O. R. Polo², M. Prieto², A. Ravanbakhsh^{3,19}, S. Sánchez-Prieto², C. E. Schlemm¹, H. Seifert¹, J. Souček²⁰, M. Steller⁸, Š. Štverák²¹, J. C. Terasa³, P. Trávníček^{10,21}, K. Tyagi^{1,22}, A. Vaivads^{4,23}, A. Vecchio^{6,24}, and M. Yedla^{3,19}.

Solar Orbiter's first Venus Flyby:

MAG observations of structures and waves associated with the induced Venusian magnetosphere

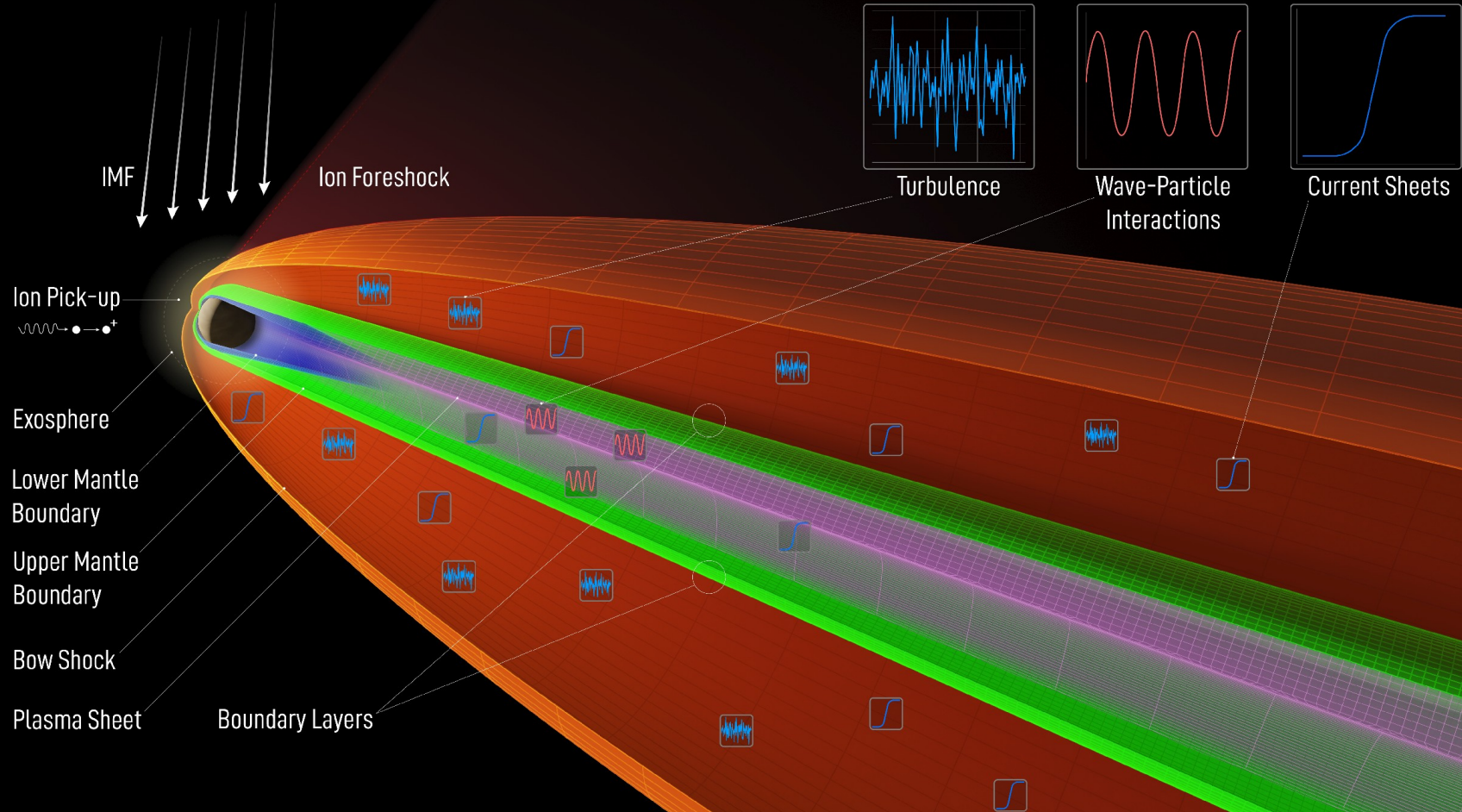
M. Volwerk¹, T. S. Horbury², L. D. Woodham², S. D. Bale³, C. Simon Wedlund¹, D. Schmid¹, R. C. Allen⁴, V. Angelini², W. Baumjohann¹, L. Berger⁸, N. J. T. Edberg⁶, V. Evans², L. Z. Hadid⁵, G. C. Ho⁴, Yu. V. Khotyaintsev⁶, W. Magnes¹, M. Maksimovic⁵, H. O'Brien², M. B. Steller¹, J. Rodriguez-Pacheco⁷, and R. F. Wimmer-Scheingruber⁸

RPW

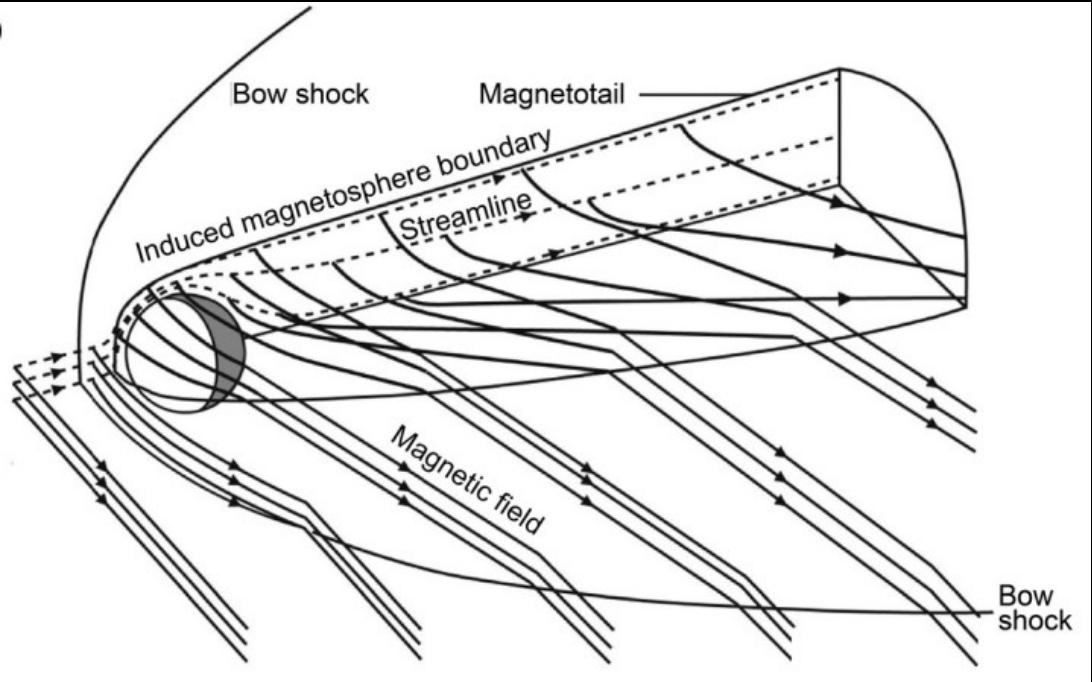
EPD/STEP

MAG

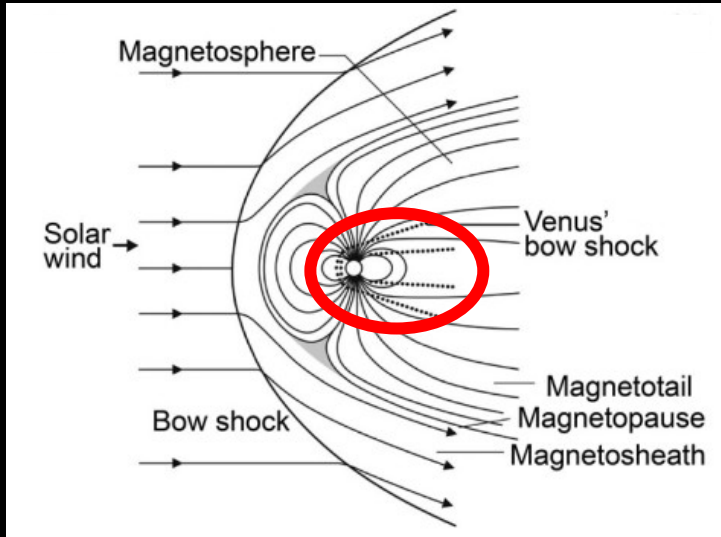
Solar Orbiter's in situ observations during the first flyby of Venus



Induced magnetosphere of Venus



Y. Futaana et al., *Space Sci. Rev.*, 2017



Mariner 10: study of the magnetic field fluctuations far downtail $\sim 70 R_V$

Pioneer Venus Orbiter: the only spacecraft that carried a plasma wave instrument (4 channel electric field receiver).

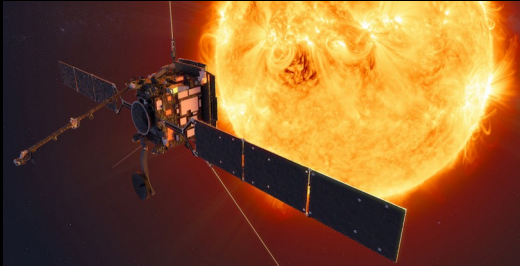
Solar Orbiter: A first opportunity to study the fields fluctuations In Venus induced magnetosphere down to $\sim 100 R_V$

Since 1960's:

- Flyby missions: Mariner 2, 5, 10; Galileo, Cassini
- Landers: Venera 4-16
- Orbiters: Pioneer Venus, Venus Express, and Akatsuki
- Atmospheric probes

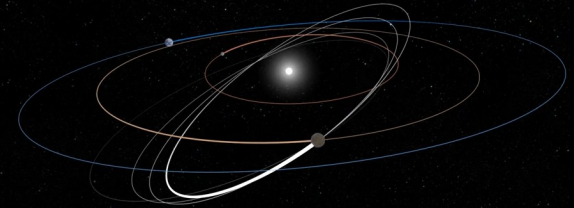
Solar Orbiter Venus flyby #1: 27/12/2020

Solar mission to explore the inner heliosphere and the polar regions of the Sun!

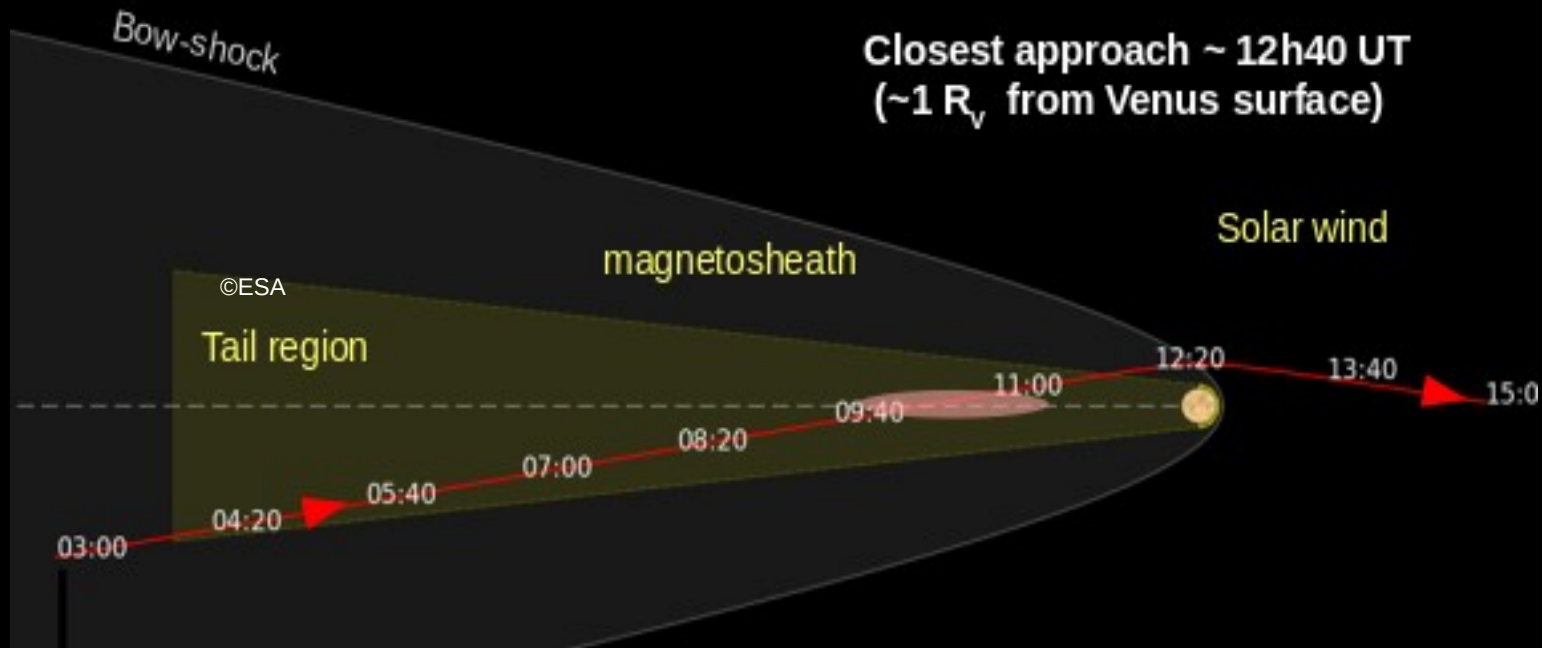


8 gravity-assists manoeuvres (7 Venus and 1 Earth)

Solar Orbiter 1st Venus flyby



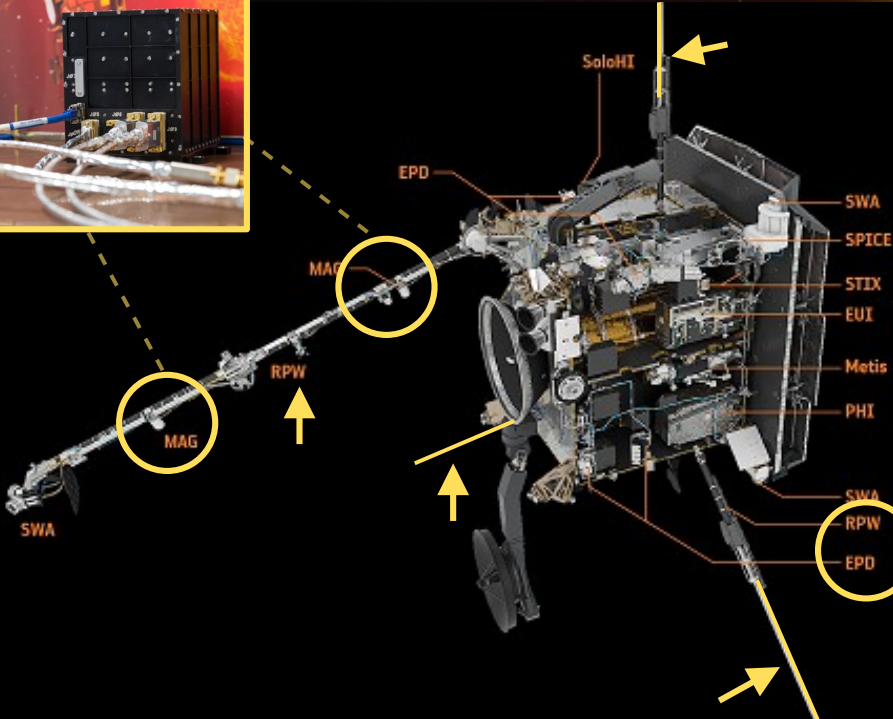
Closest approach ~ 12h40 UT
(~1 R_V from Venus surface)



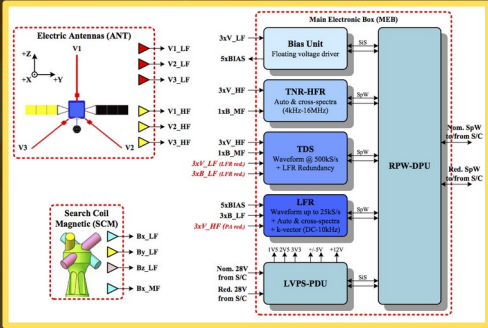
Operational in-situ instruments: MAG, RPW and EPD/STEP



Magnetometer/MAG
(PI: T. S. Horbury)

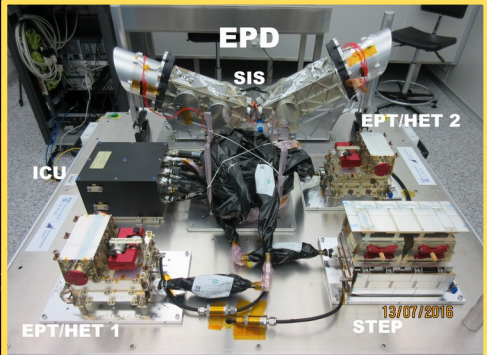


Radio and Plasma Wave consortium:
Electric field antenna, Search-Coil magnetometer
& the main electronic box (PI: M. Maksimovic)



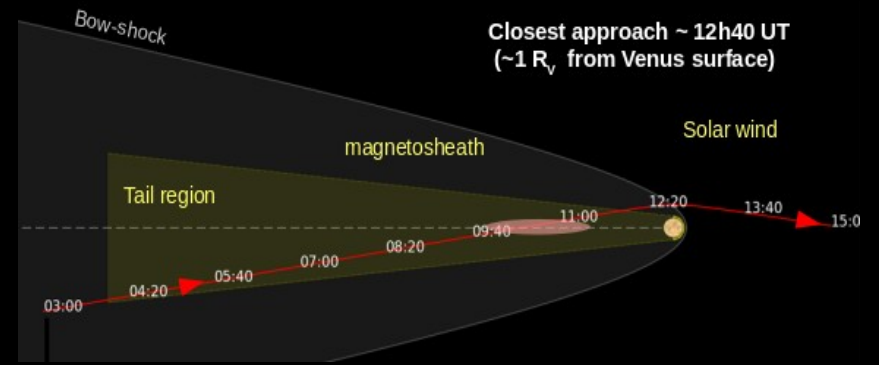
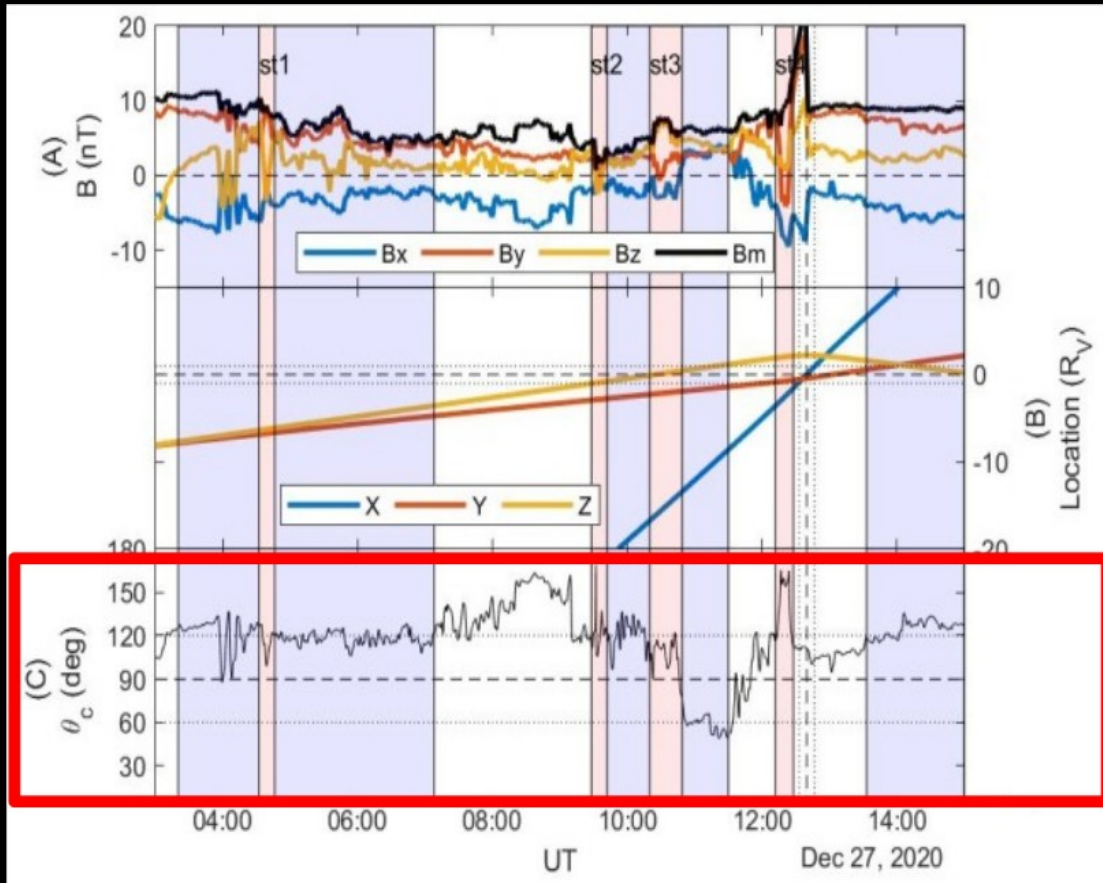
E: ~DC - 16 MHz
B: 1Hz -1 MHz

Energetic Particle Dectector (EPD) /
SupraThermal Electrons and Protons (STEP)
(PI: J. Pacheco)



STEP:
Ions: 4-80 keV

Structure of Venus' induced magnetosphere



Cone angle

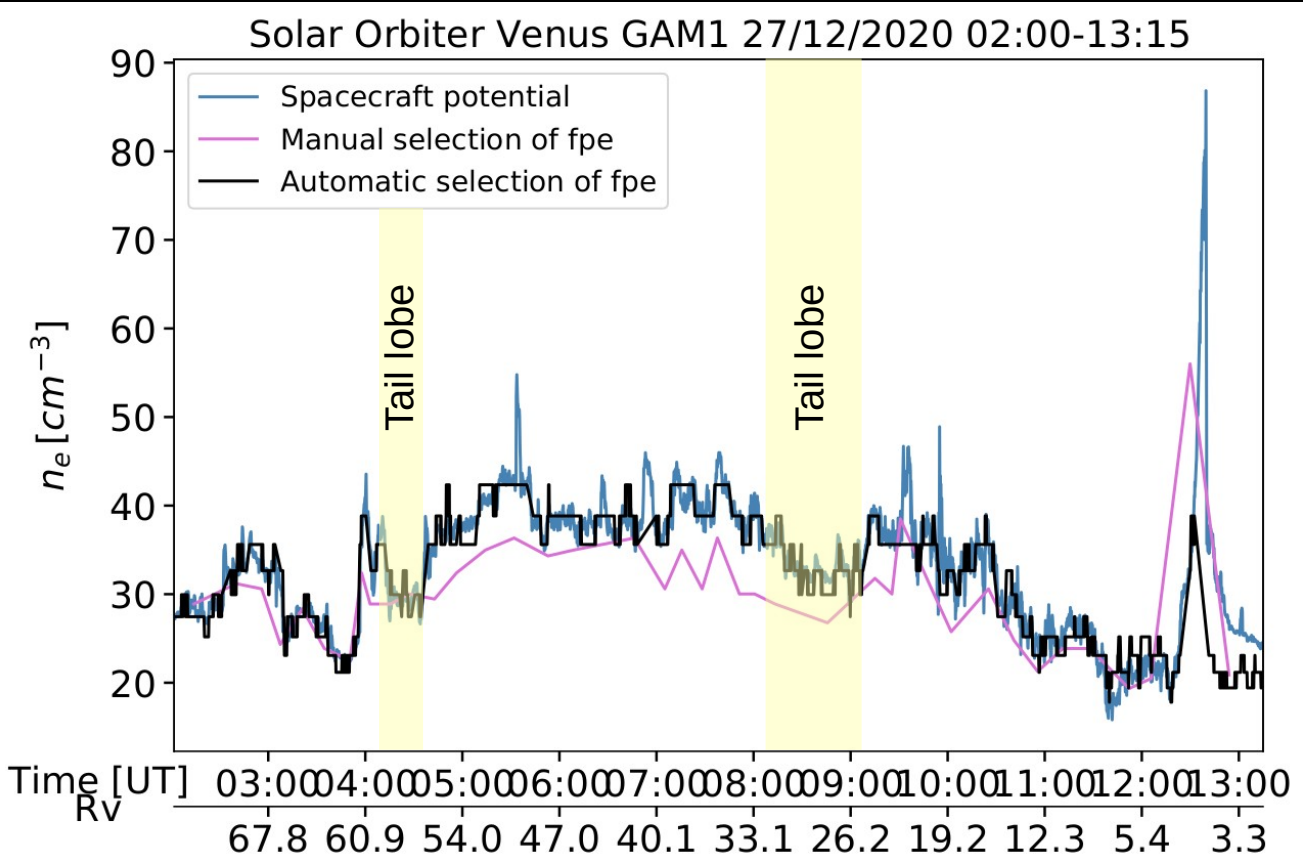
$\Theta_c = 0^\circ / 180^\circ \rightarrow$ sunward/anti-sunward

$\Theta_c = 90^\circ \rightarrow$ perpendicular to the Venus-Sun line

$$\theta_c = \tan^{-1} \left(\frac{\sqrt{B_y^2 + B_z^2}}{B_x} \right)$$

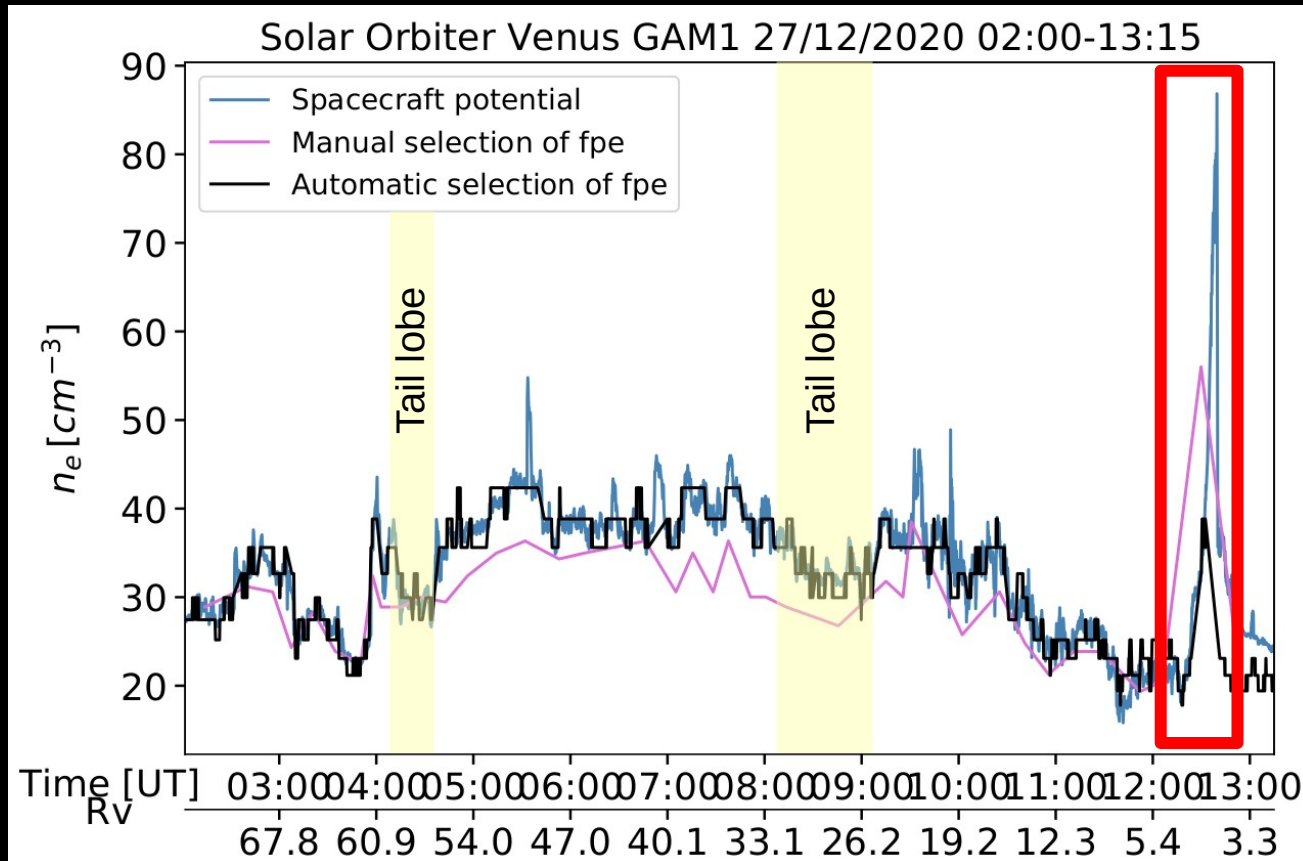
Adapted from
M. Volwerk et al., A&A, 2021

Large structure of the induced magnetosphere

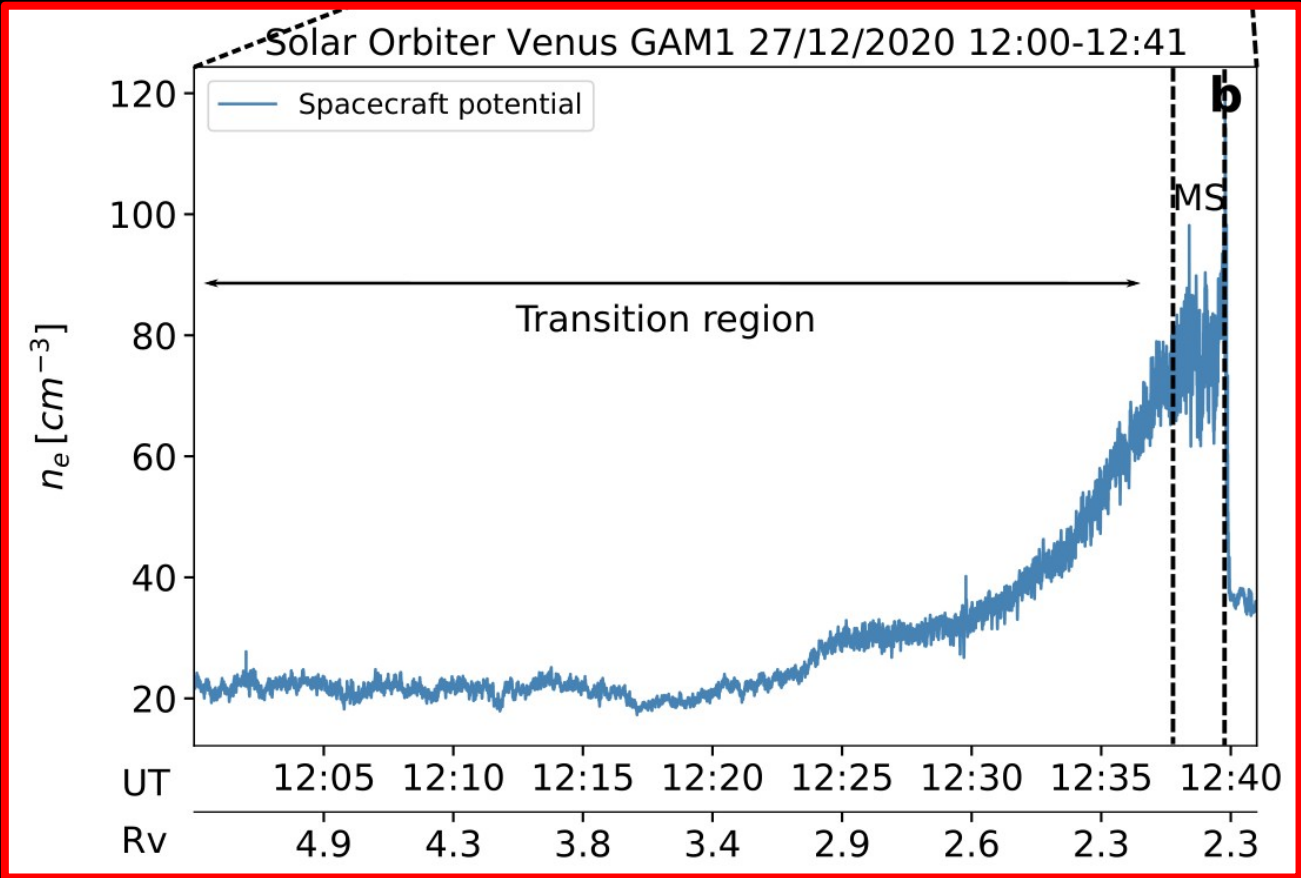


L. Z. Hadid et al., A&A, 2021

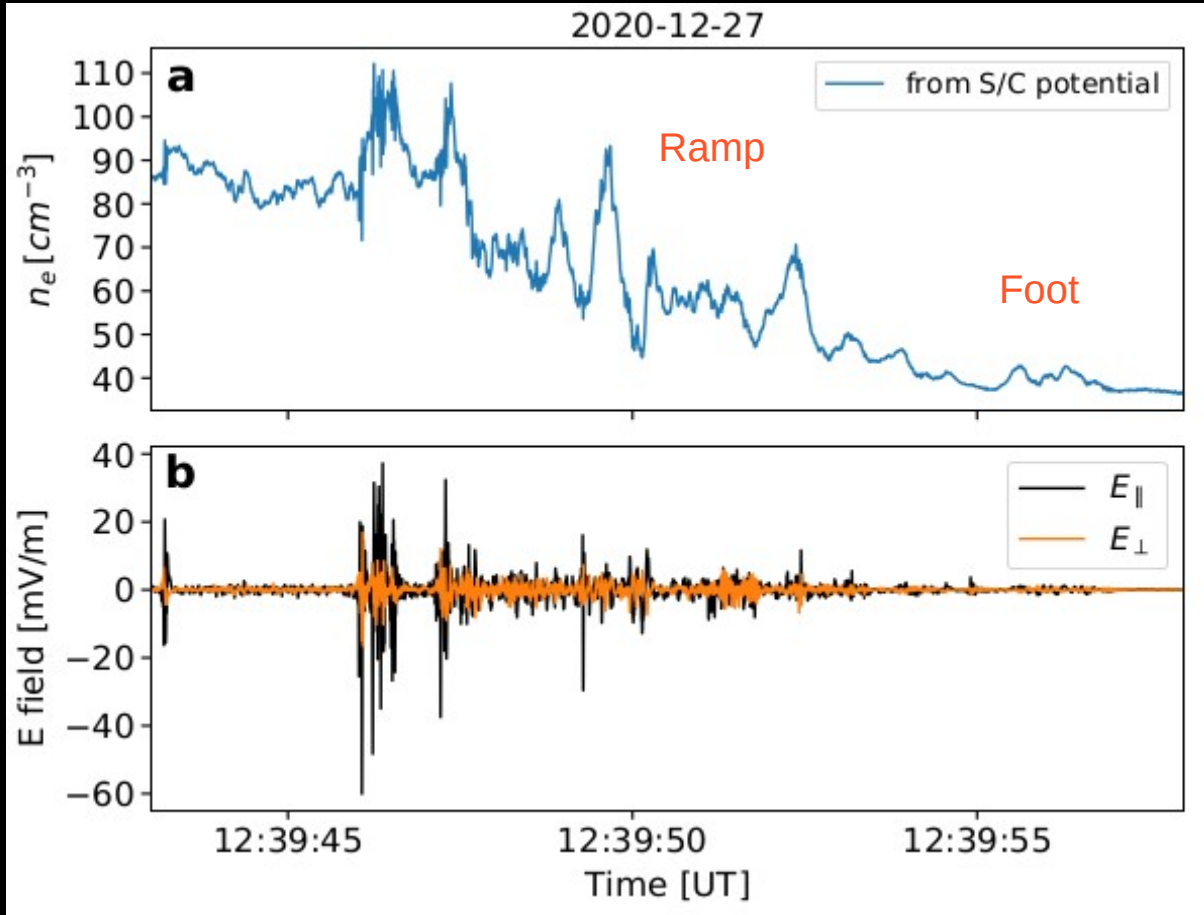
Large structure of the induced magnetosphere



Large structure of the induced magnetosphere



Bow-shock structures and waves

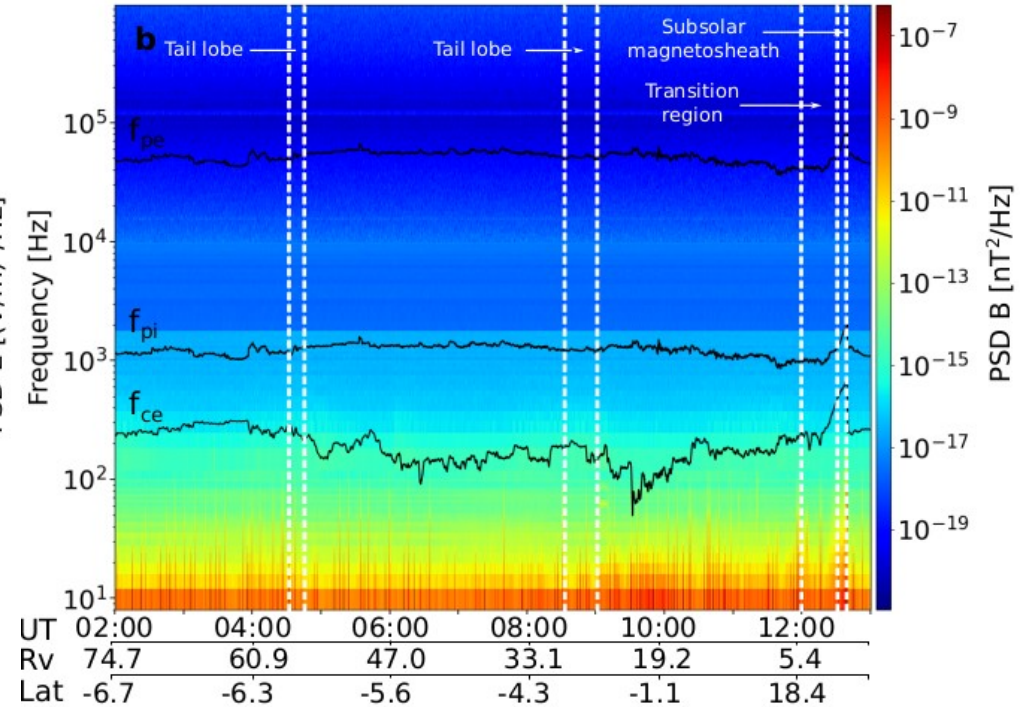
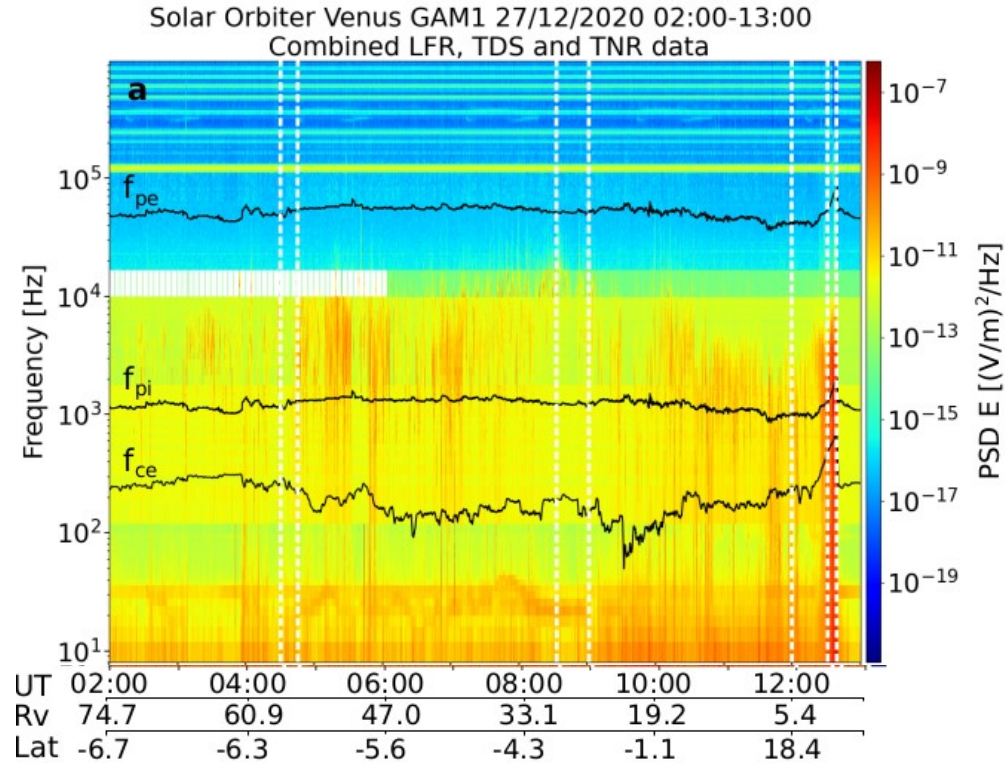


Presence of “shock-ripples”
[Gingell et al. 2017, Johlander et al. 2018] or
large amplitude whistler waves
[Wilson et al. 2017]

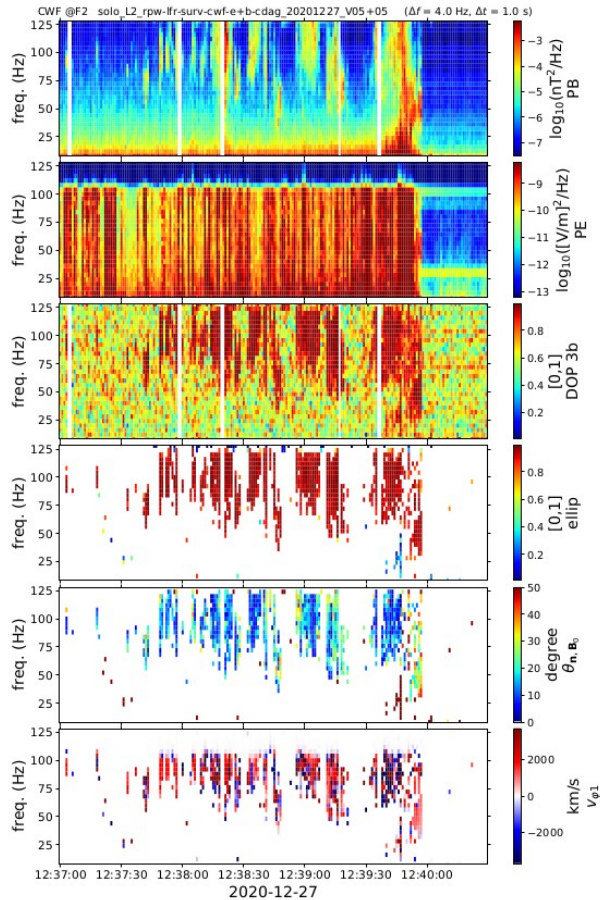
L.Z. Hadid et al. A&A, 2021

A. Dimmock et al., A&A, 2021, submitted

Magnetic and electric field power spectra from the RPW instrument



Whistler waves in the magnetosheath



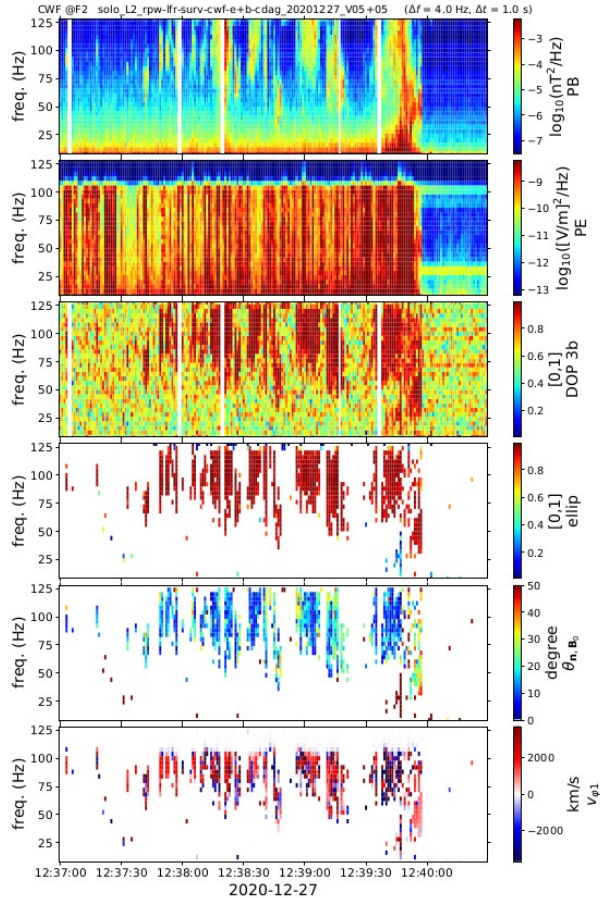
Dominance of electromagnetic, circularly polarized whistler waves

Degree of Polarization (DOP) ~ 1
→ Presence of a single plane wave

Ellipticity ~ 1
→ Circularly polarized wave

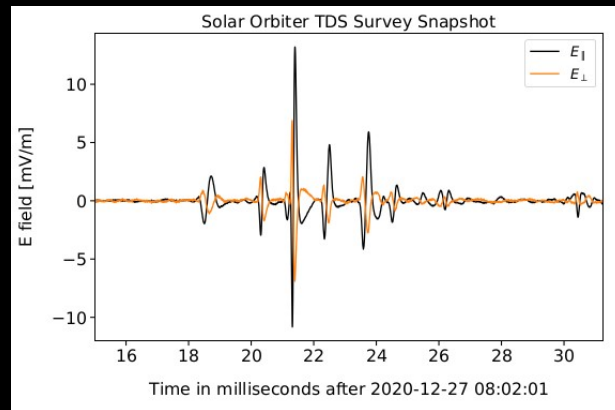
$\Theta_{n, B_0} \sim 15^\circ$
→ wave normal vector is parallel to B_0
 $\Theta_{n, B_0} < 90^\circ$
→ n and B_0 point in the same direction → right hand polarized wave

High frequency electrostatic waves in the magnetosheath

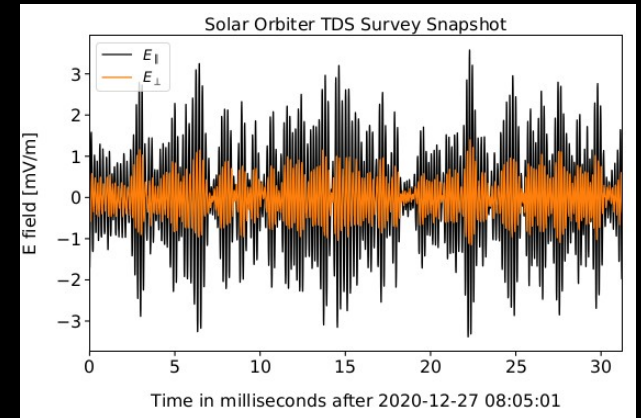


Dominance of electromagnetic, circularly polarized whistler waves

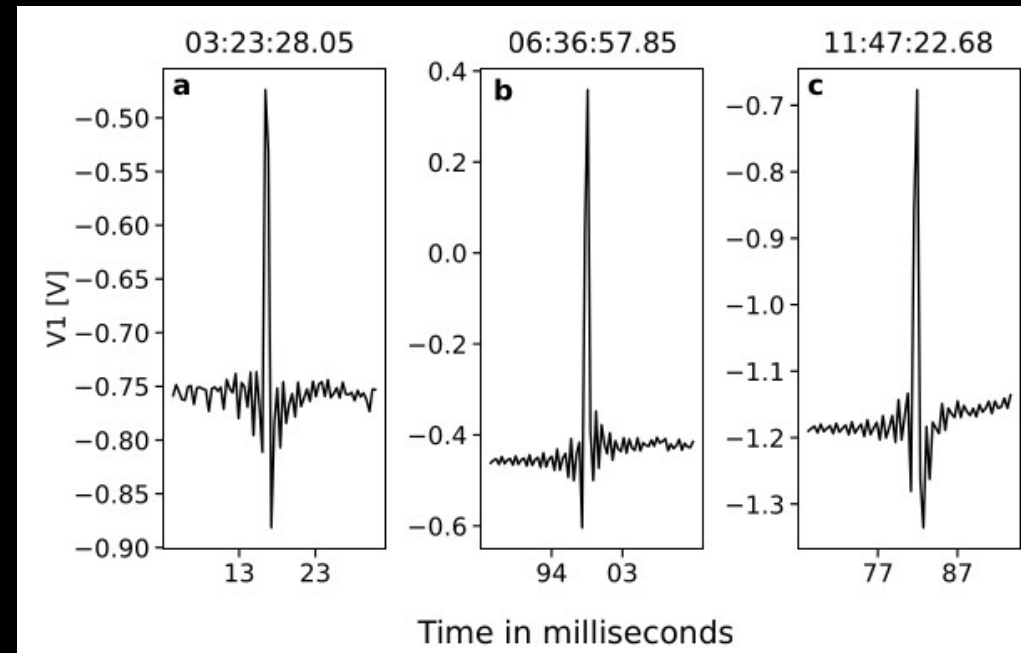
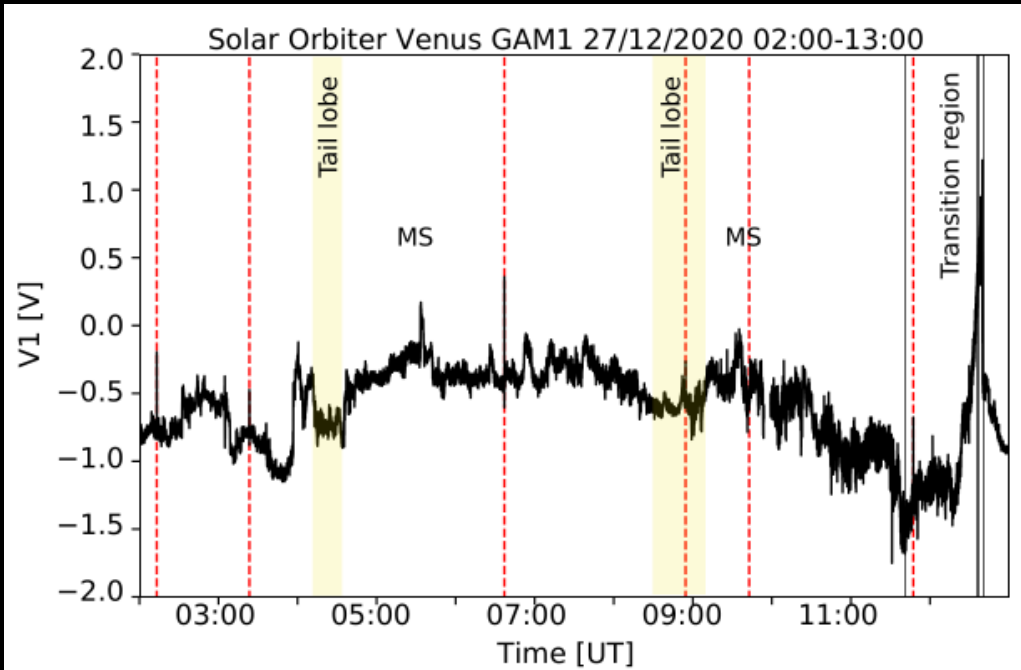
Solitary waves (~2 kHz)



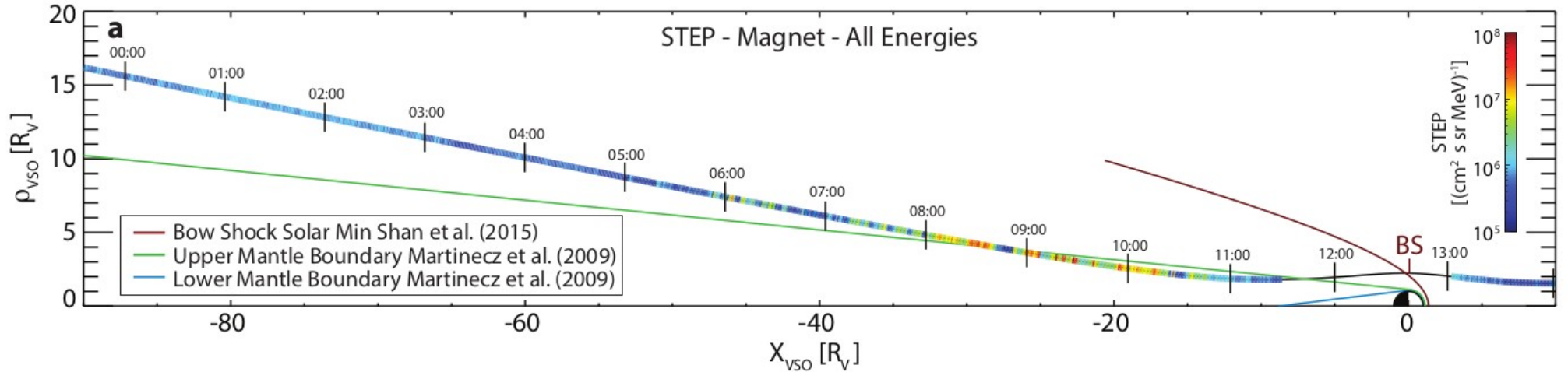
Electrostatic waves (Ion acoustic waves ~10 kHz)



Dust impacts



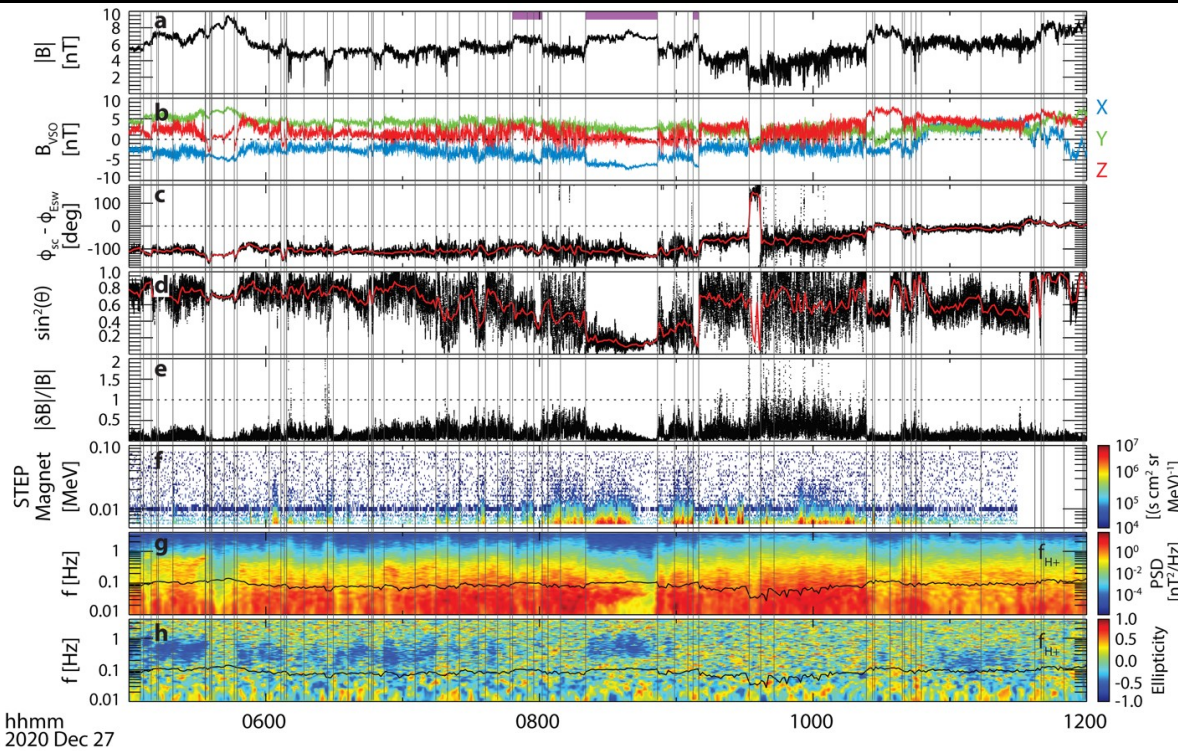
Bursty energetic ion enhancements in the magnetotail and foreshock



R. Allen et al., A&A, 2021

Enhancement of the suprathermal flux (~ 30 keV) near the empirical upper mantle boundary and the foreshock-region

Main processes leading to particles acceleration

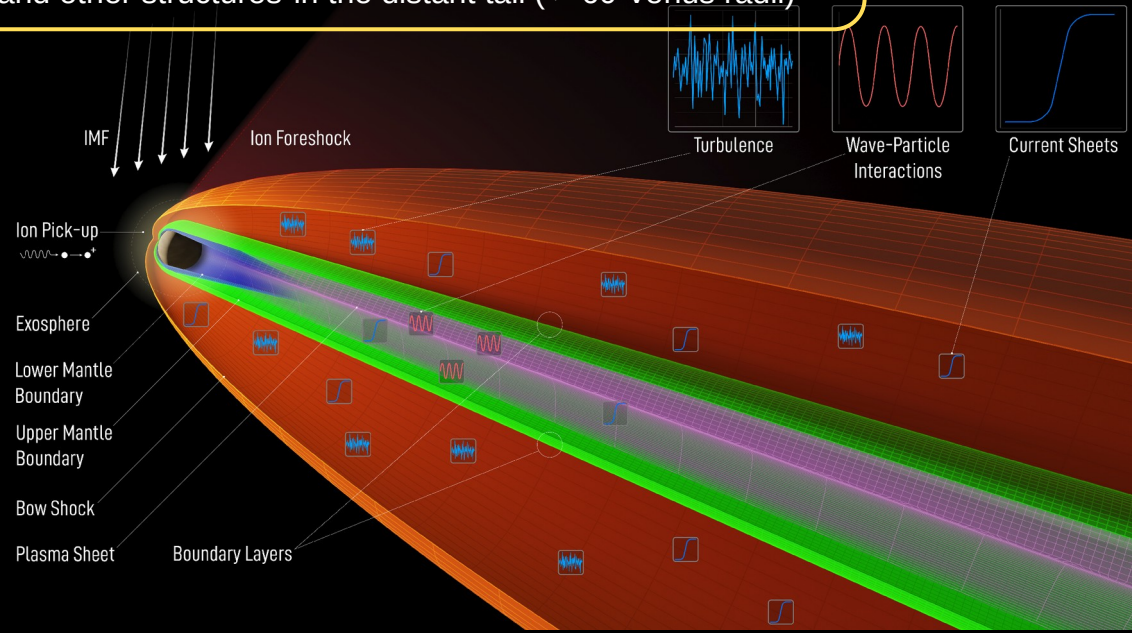


- 1) Convective solar wind electric field
- 2) Pick-up ions processes
- 3) Turbulence
- 4) Wave-particle interactions (ion-cyclotron waves and high frequency electrostatic waves)
- 5) Boundary layer heating
- 6) Current sheet acceleration

Solar Orbiter's in situ observations during the first flyby of Venus

1. Large scale magnetic field structure

- Boundary crossings: bowshock, ion composition boundary, plasmashet
- Bow-shock geometry: clear quasi-perpendicular bow-shock
- Draping of the magnetic field over North pole
- Evidence of plasmoids/fluxrops, magnetic reconnection site and other structures in the distant tail (> 60 Venus radii)



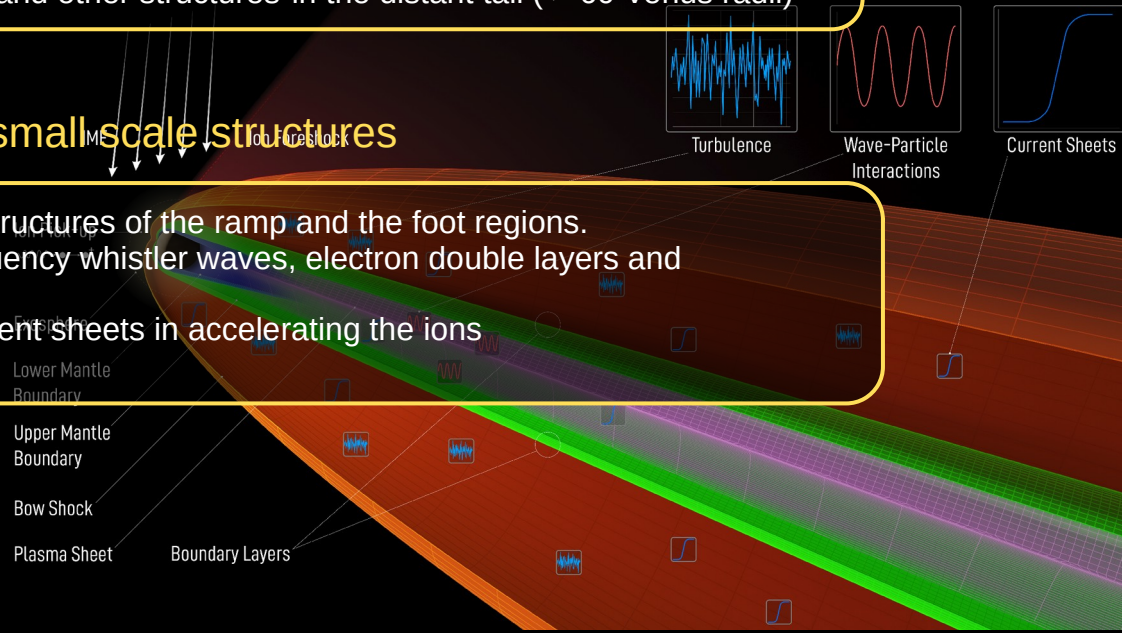
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2. Wave properties, acceleration processes and small scale structures

- Bow-shock: whistler waves and electric and magnetic sub-structures of the ramp and the foot regions.
- Magnetosheath: presence of ion cyclotron waves, high frequency whistler waves, electron double layers and electrostatic waves.
- Potential role of turbulence, ICWs, boundary layers and current sheets in accelerating the ions
- Dust impacts



Solar Orbiter's *Astronomy & Astrophysics* special issue and more...

Overview papers

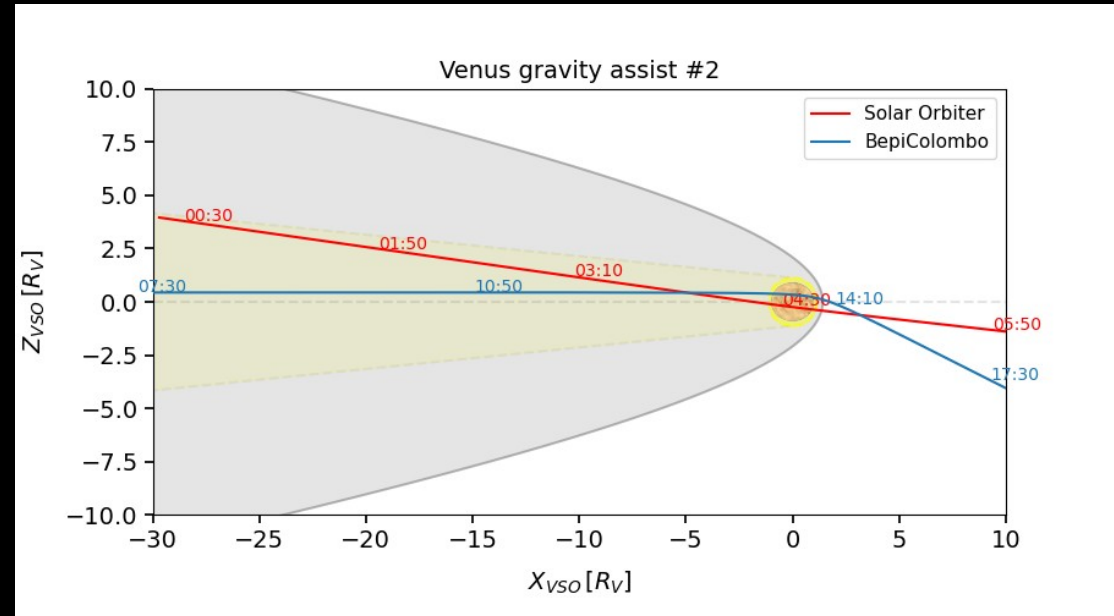
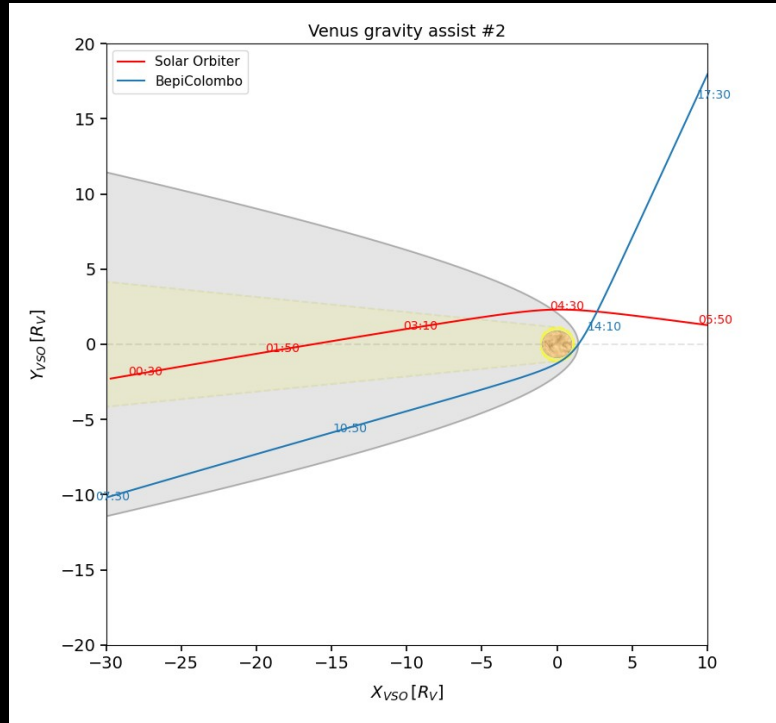
1. MAG: "Solar Orbiter Venus 1 flyby: Magnetometer observation", **M. Volwerk et al.**, accepted
2. EPD: "Energetic Ions in the Venusian System: Insights from the First Solar Orbiter Flyby", **R. Allen et al.**, accepted
3. RPW: "Solar Orbiter's first Venus flyby: observations from the Radio and Plasma Wave instrument", **L. Z. Hadid et al.**, accepted

Detailed studies

1. Structures at bow-shock of Venus, **A. Dimmock et al.**, submitted
2. Turbulence and intermittency in Venus induced magnetosphere, **L. Sorriso-Valvo et al.**, in prep
3. Comparisons between the electron number densities and simulations of SW-Venus interaction, **K. Stergiopoulou et al.**, in prep
4. Dust impacts statistics, **M. Morooka et al.** In prep



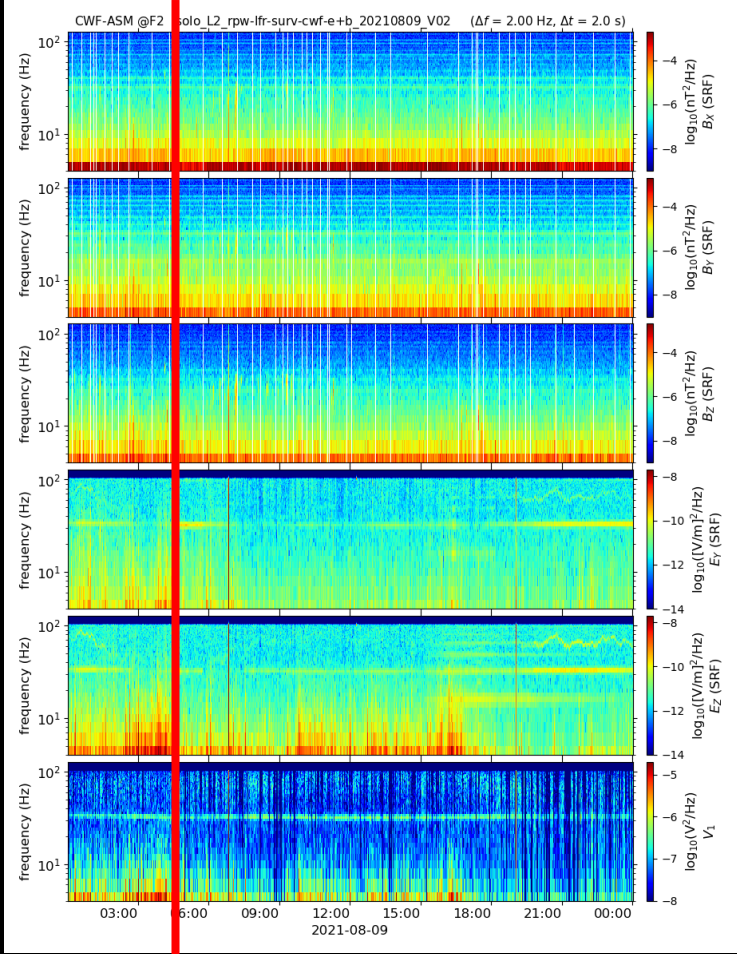
Solar Orbiter's second Venus flyby: August 09, 2021



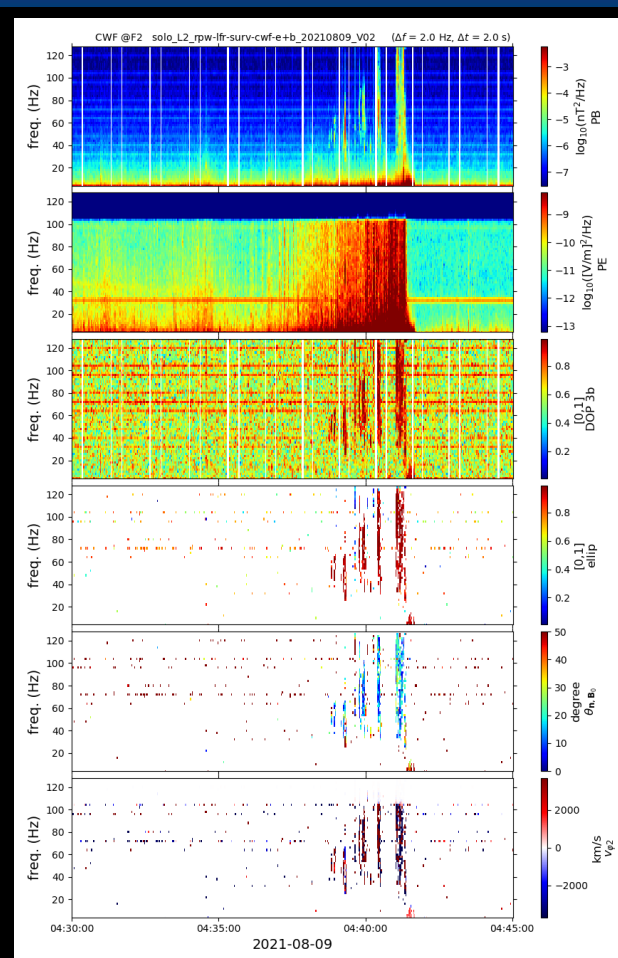
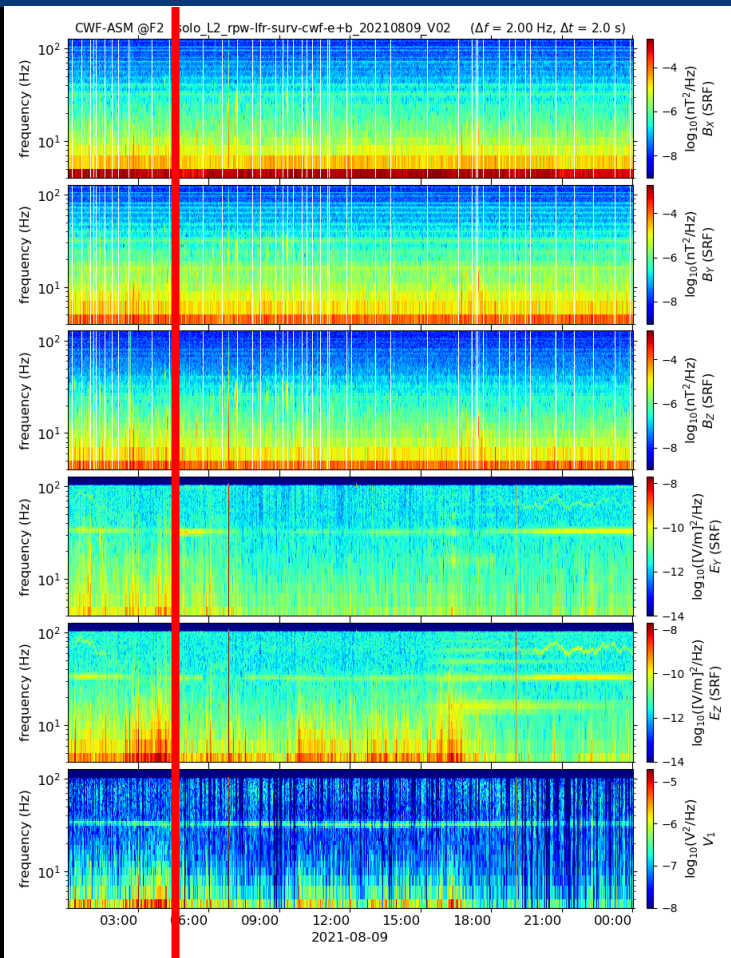
CA SoIo ~ 2.3 R_V (2021-08-09 ~ 04:43)

CA Bepi ~ 1.1 R_V (2021-08-10 ~ 13:51)

Solar Orbiter's VGAM2: preliminary LFR data

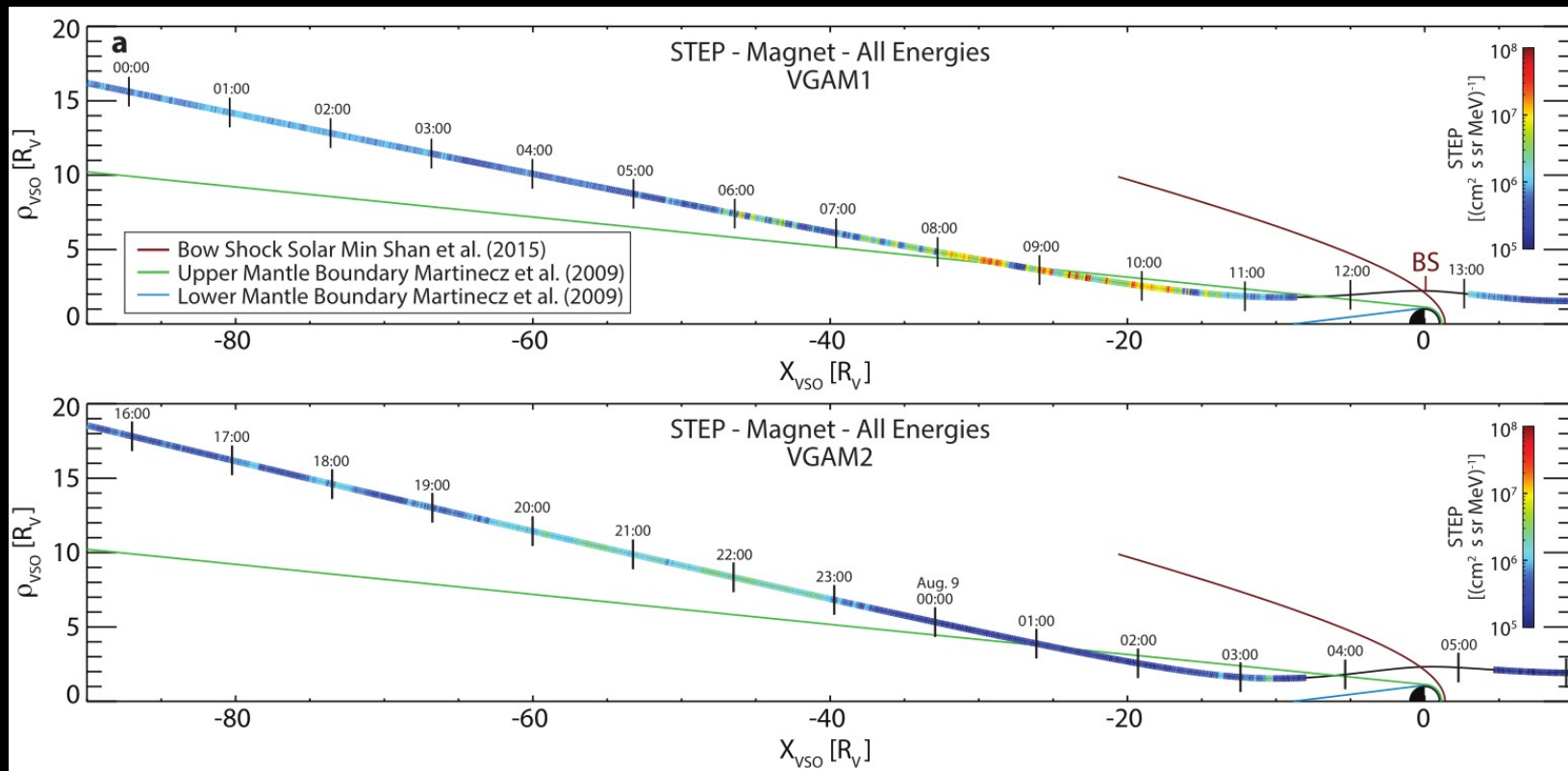


Solar Orbiter's VGAM2: preliminary LFR data



Solar Orbiter's VGAM2: preliminary EPD data

VGAM1



VGAM2

A detailed illustration of a satellite in space. The satellite is a complex, dark-colored structure with various instruments, antennas, and solar panels. It is positioned in the center of the frame, with a large, bright, orange-yellow sun in the background. The sun's surface is highly textured, showing solar flares and other activity. The background is a deep red-orange color, suggesting a close proximity to the sun. The overall scene is set against a dark, starry space background.

Thank you!